

Fluorescence Efficiency and Visible Re-emission Spectrum of Tetraphenyl Butadiene Films at Extreme Ultraviolet Wavelengths



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Instrumentation Brown Bag
Seminar**

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arXiv:1104.3259

Outline:

- Noble gas detectors and scintillation light
- Wavelength shifting films (TPB), and previous work
- Our work:
 - Experimental apparatus and systematic checks
 - Visible re-emission spectrum
 - Absolute fluorescence efficiency
- Plans for future work at LBL
- Conclusions

Noble Gas Detectors and Scintillation Light

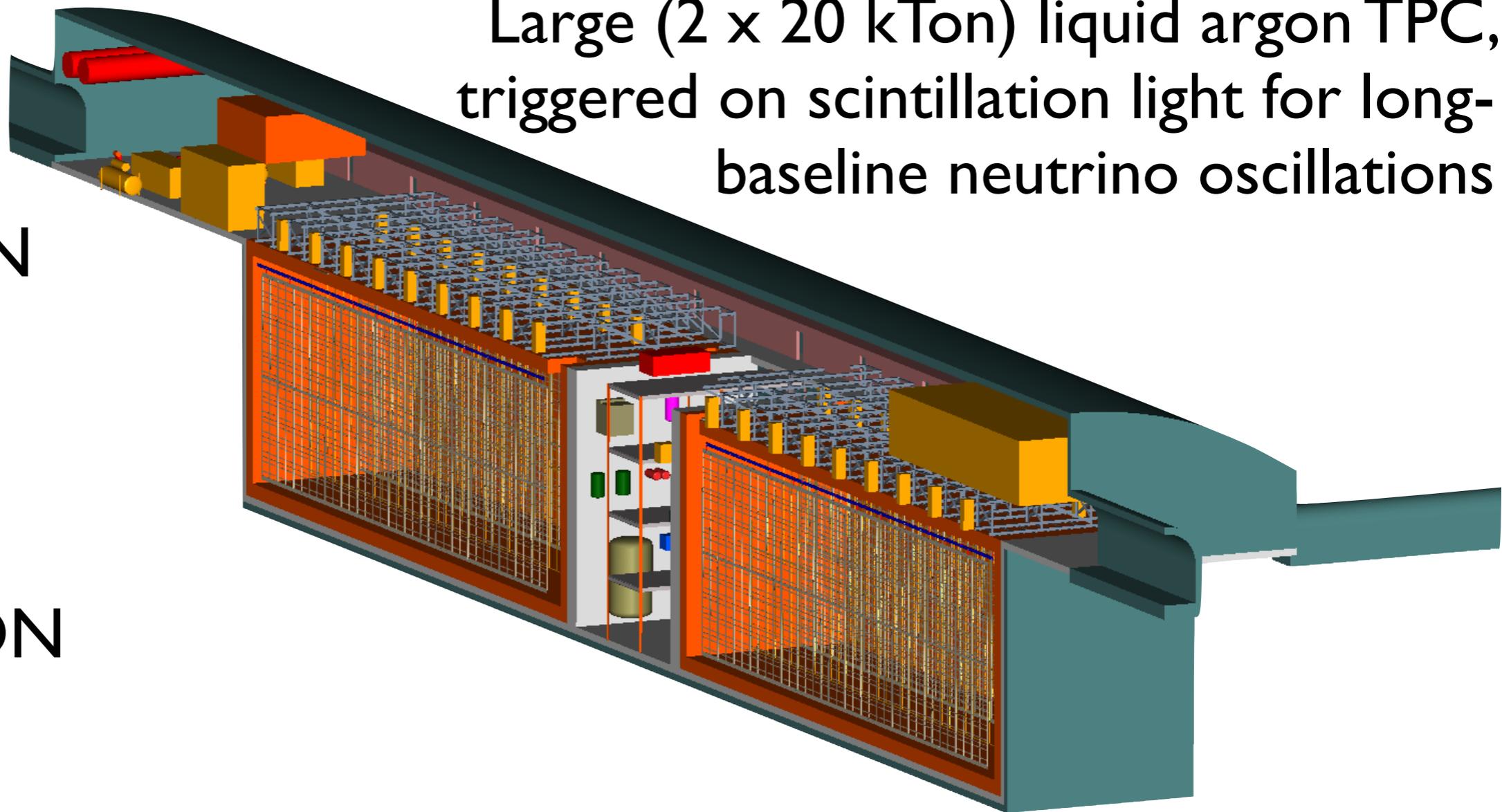
Handling 100-nm or so photons is a pain in the neck!

Noble Gas Detectors

- Long history of use in proportional counters and TPCs
- Quite a few proposed ones that will also collect scintillation light:

- LBNE
- NEXT
- CLEAN
- nEDM
- LUX
- XENON

Large (2 x 20 kTon) liquid argon TPC, triggered on scintillation light for long-baseline neutrino oscillations

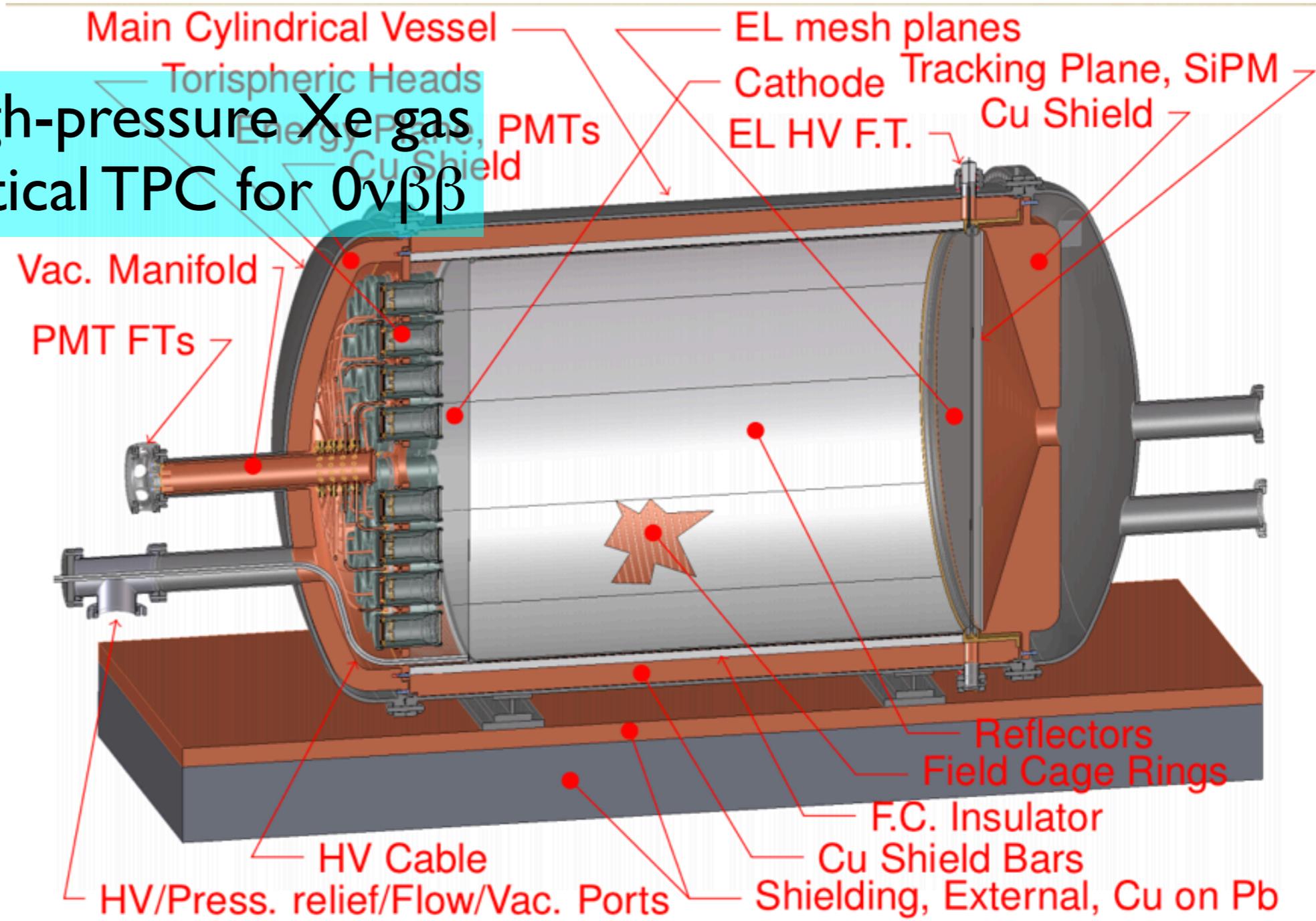


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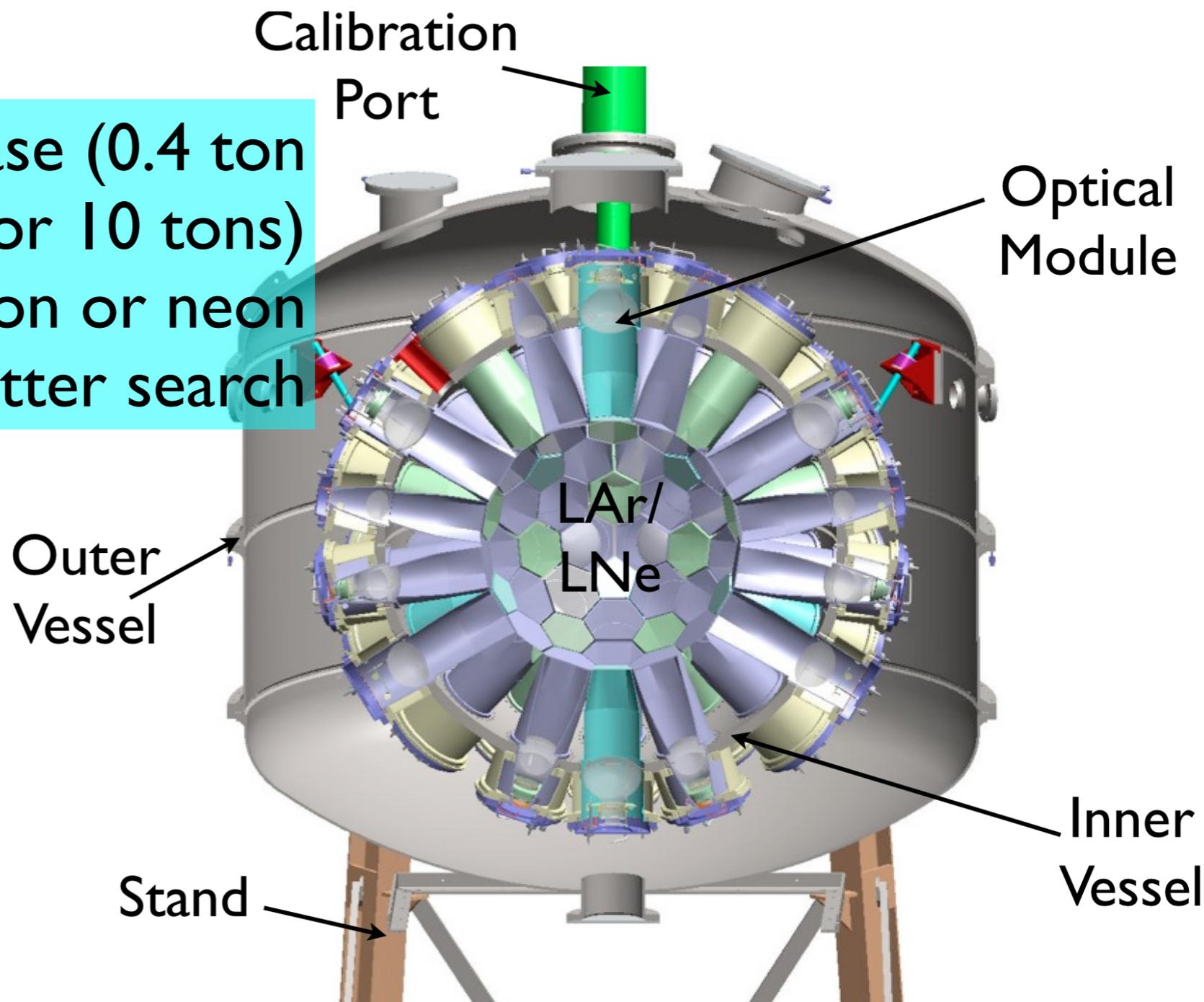
High-pressure Xe gas optical TPC for $0\nu\beta\beta$



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- Single-phase (0.4 ton now, plans for 10 tons)
liquid argon or neon
dark matter search

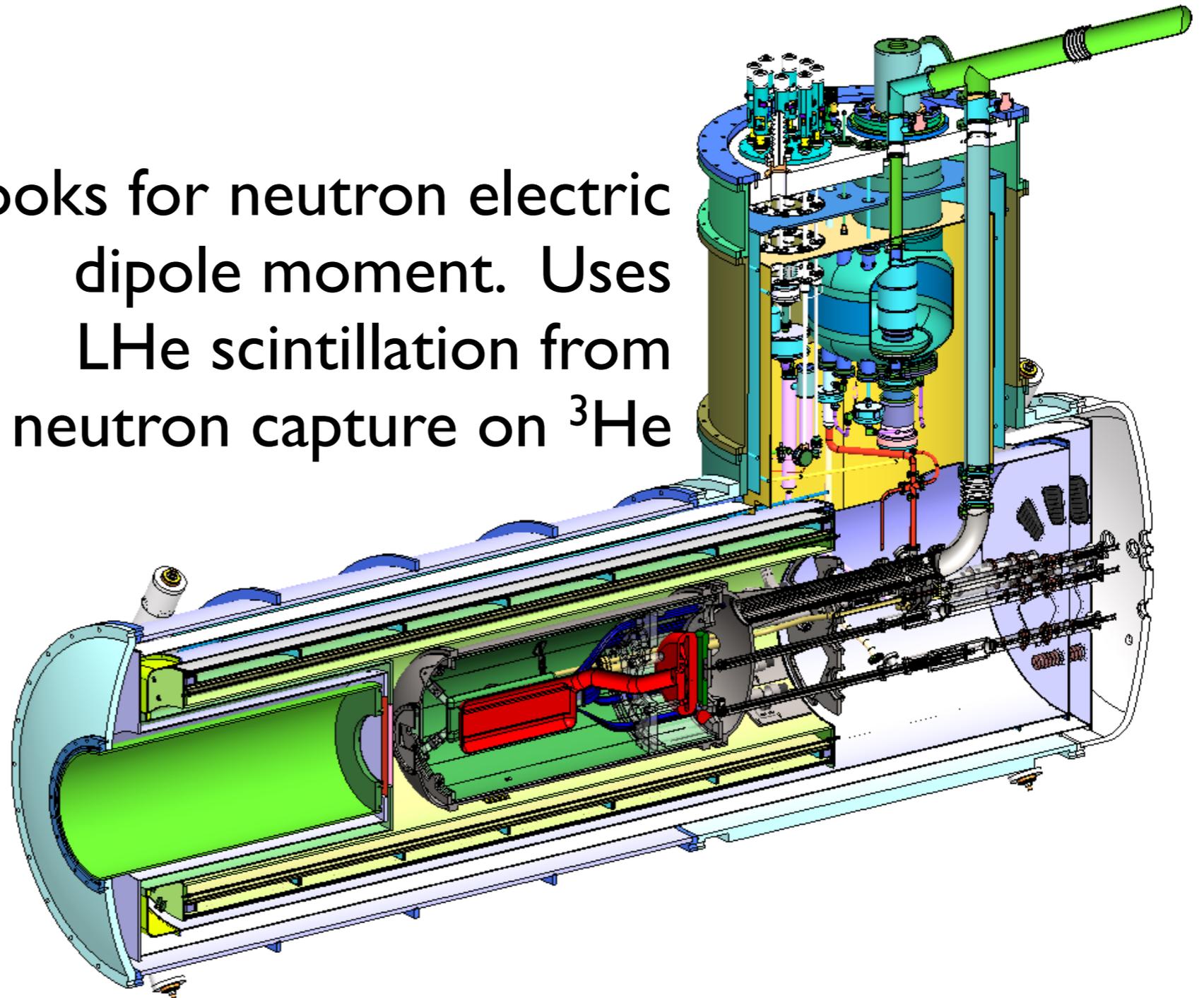


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Looks for neutron electric dipole moment. Uses LHe scintillation from neutron capture on ^3He

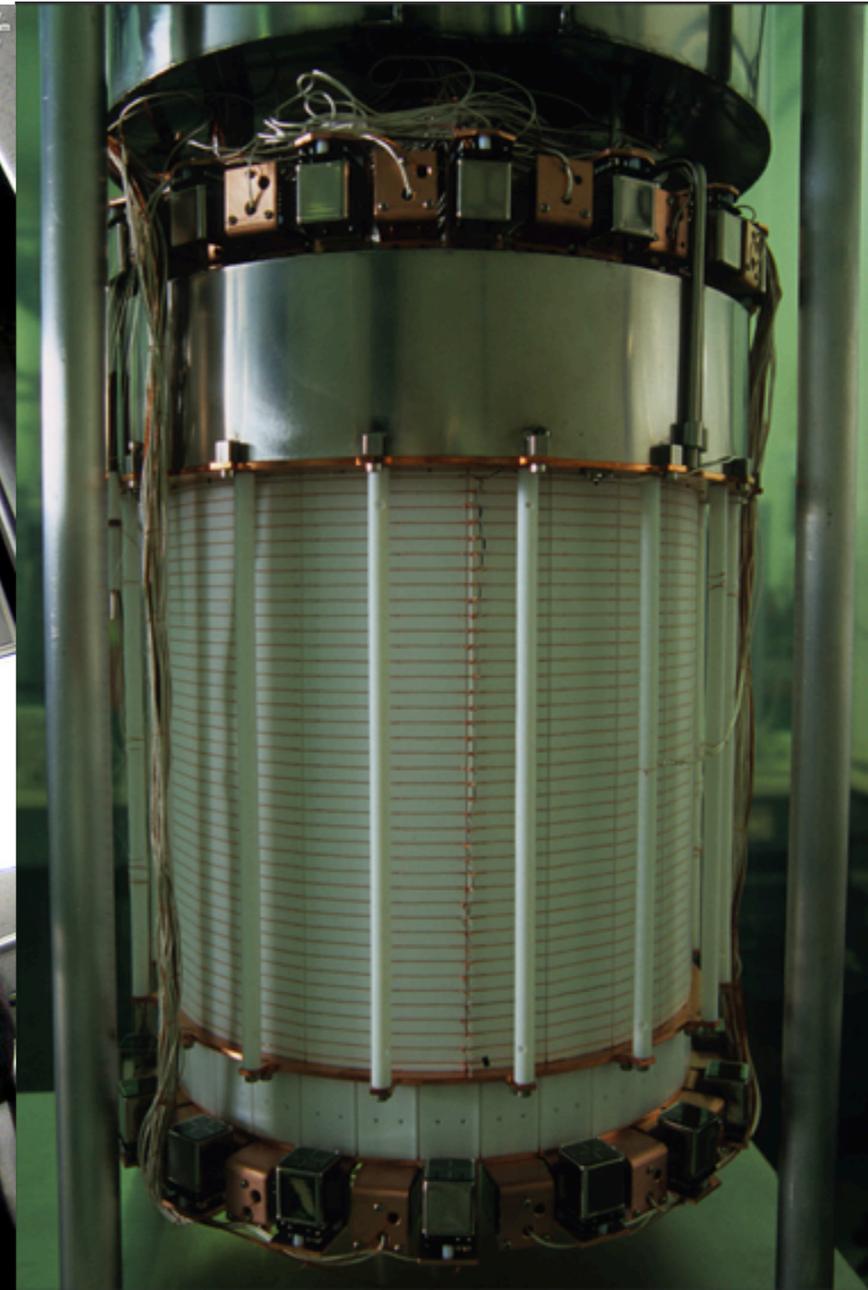
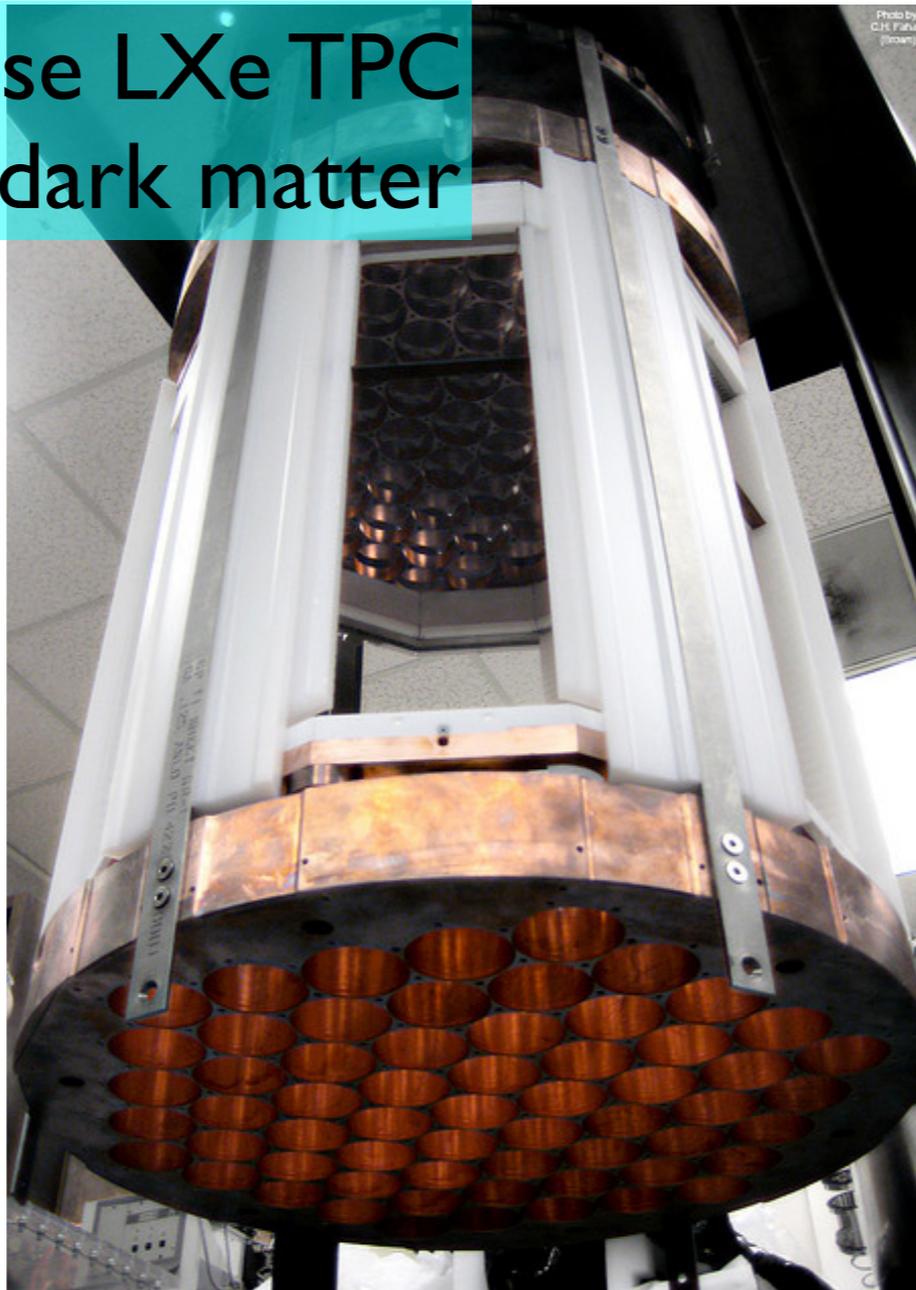


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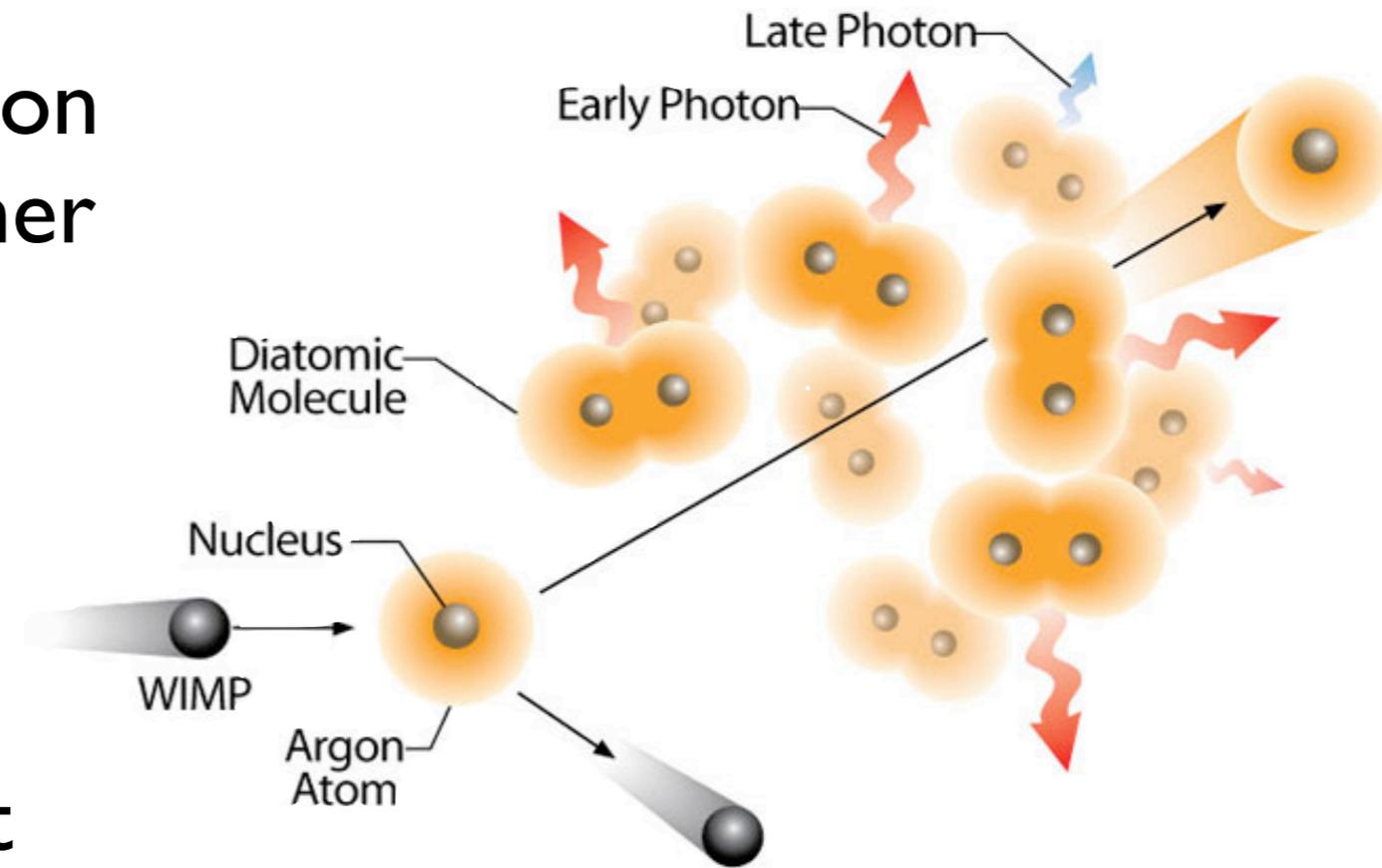
- LBNE
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Dual-phase LXe TPC
for dark matter



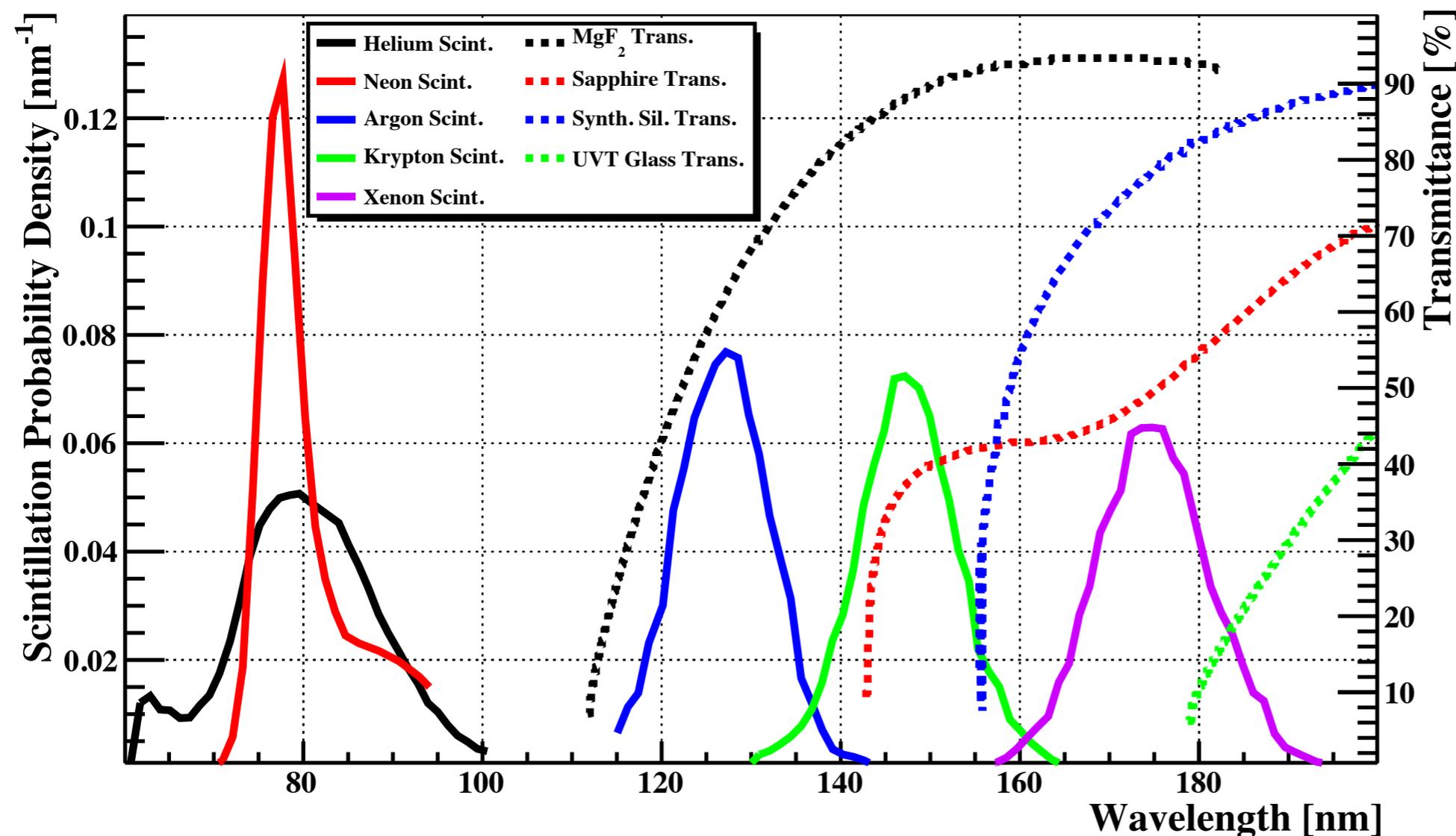
Noble Gas Scintillation

- Noble gases create scintillation light differently than most other materials
- Scintillation light comes from the breakup of dimers where one partner atom is either excited or ionized by incident radiation
- This leads to two important consequences:
 - Near perfect transparency to its own scintillation light
 - Two scintillation time constants: one from dimer breakup involving excited atoms (prompt light, few ns) and one from ionized atoms (later light, tens of ns to few μ s)



Here's the problem...

- The scintillation light is well into the extreme ultraviolet!
- Short enough wavelength that everything interacts with them, but not energetic enough to penetrate like x rays



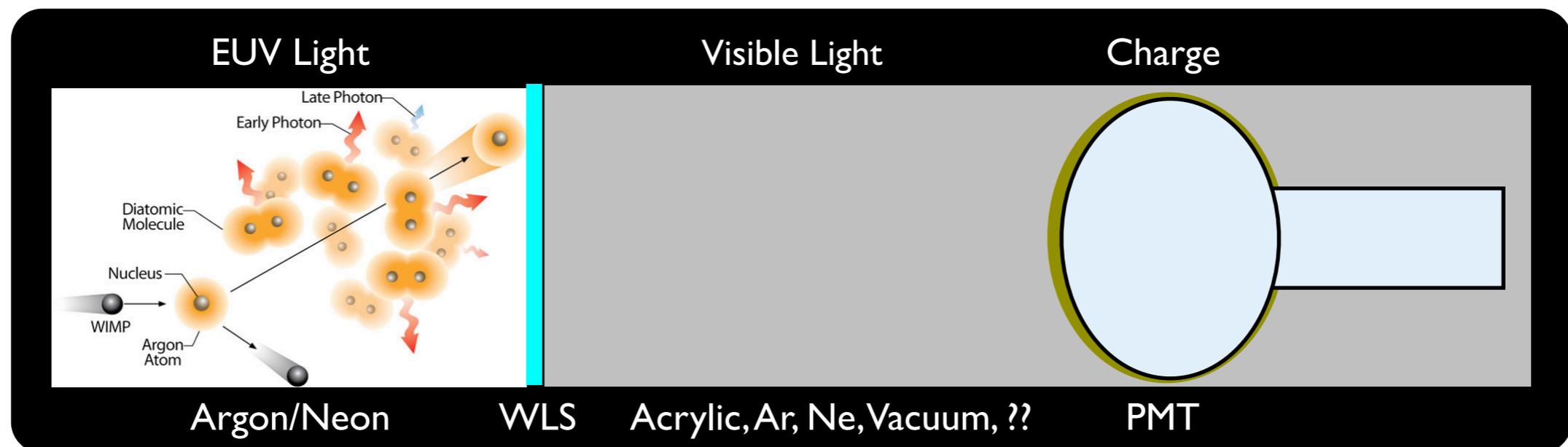
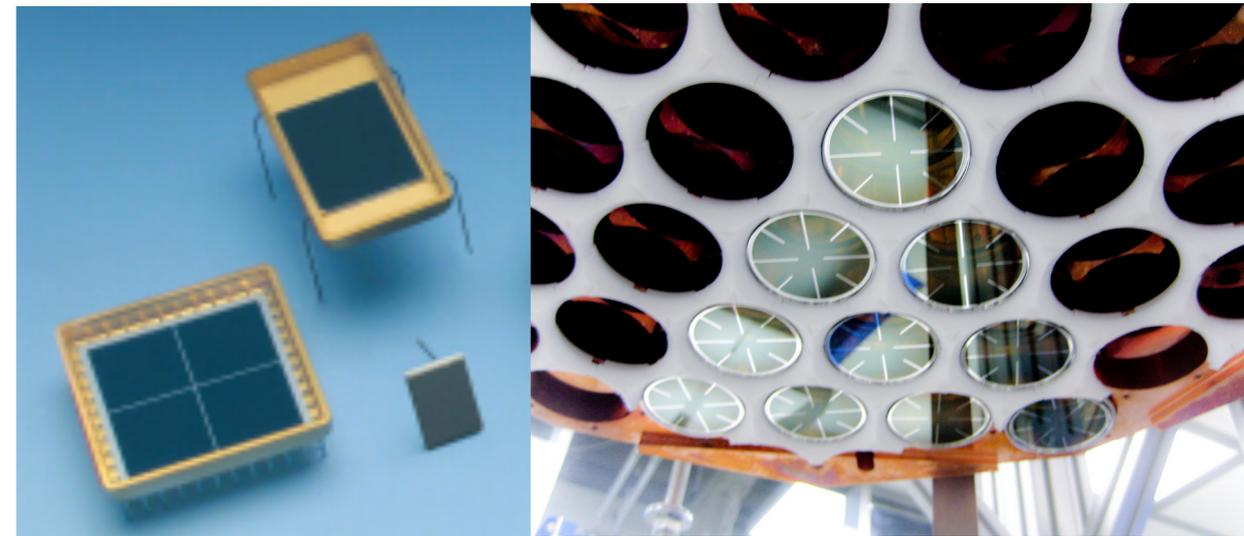
**Windows
stop
working!**

Wavelength Shifting Films and Previous Work

Why has no one done this before???

What to do with these troublesome photons?

- Some devices are directly sensitive to them
- Solid state devices can be sensitive down to below 100 nm. Small area, often slow
- Some PMTs sensitive down to 160 nm, UV-transmitting window limits area
- Usually, you need a wavelength shifting film:



What to do with these troublesome photons?

EUV Light

Visible Light

Charge



Argon/Neon

WLS

Acrylic, Ar, Ne, Vacuum, ??

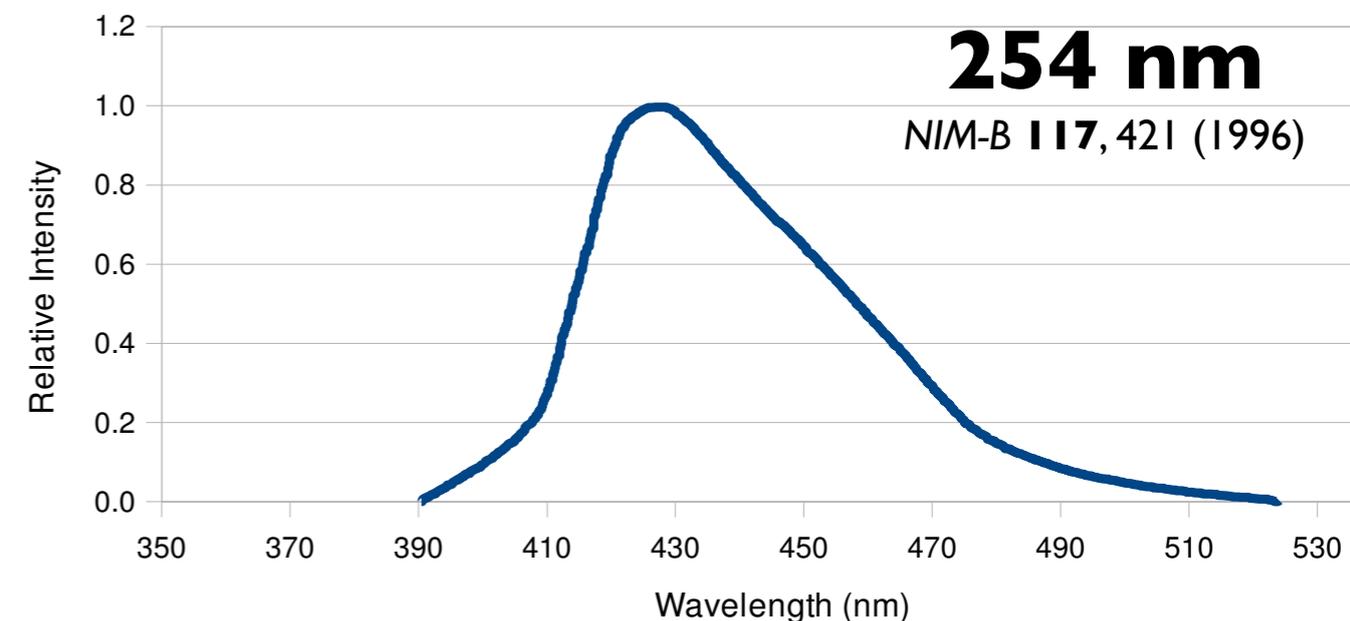
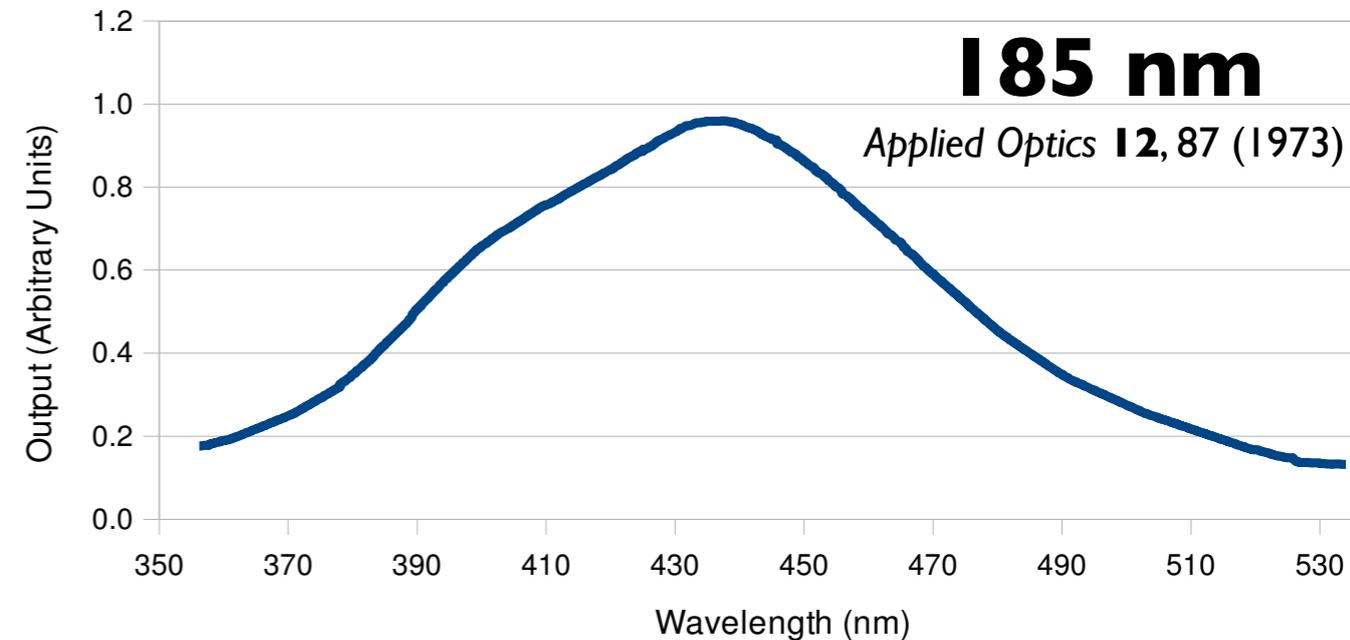
PMT

When you do this, you get “easier to detect” photons, but you give up all of your direct light

Any analysis that requires detailed understanding of your optical train becomes a bit more complicated

Hasn't Someone Done This Already?

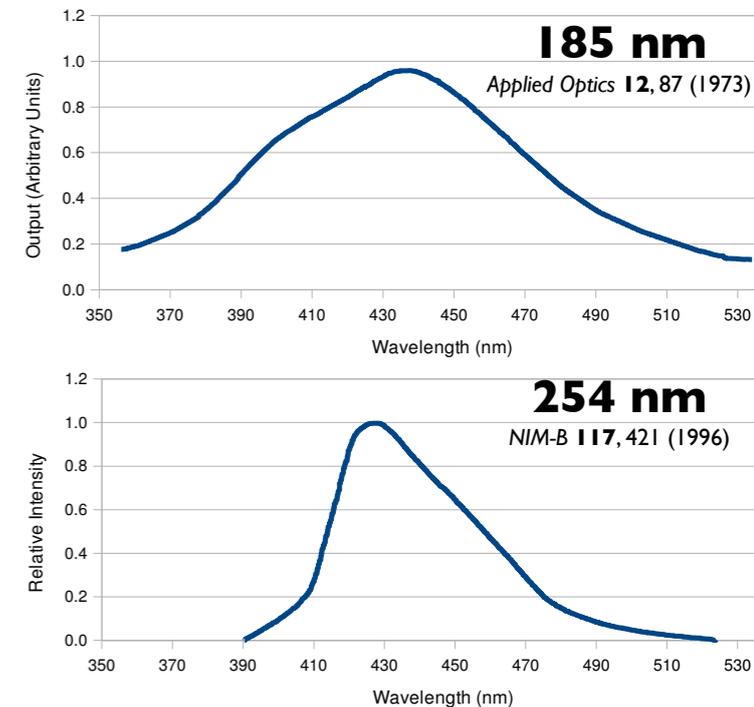
- Yes, but...
- There was a lot of ambiguity in the shape of the re-emission spectrum as a function of input wavelength
- Previous efficiency measurements were made relative to other fluors whose absolute efficiency was uncertain to about a factor of two!



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Table reproduced from "Techniques of Vacuum Ultraviolet Spectroscopy," James A. R. Samson, ©1967



Absolute Efficiency [%]			Layer Thickness [mg / cm ²]
2537 Å	1216 Å	304 Å	
99	94	41	2-4
65	62-80		5
50	38	2	
64		1-2	
25		6	
60		?	
			2 mm ^a

^aSample was a plaque pressed 2 mm thick

Our TPB Study

How I spent a good chunk of last year...

Experimental Apparatus

Since we are observing individual photons, we care about the efficiency as a ratio of photon rates.

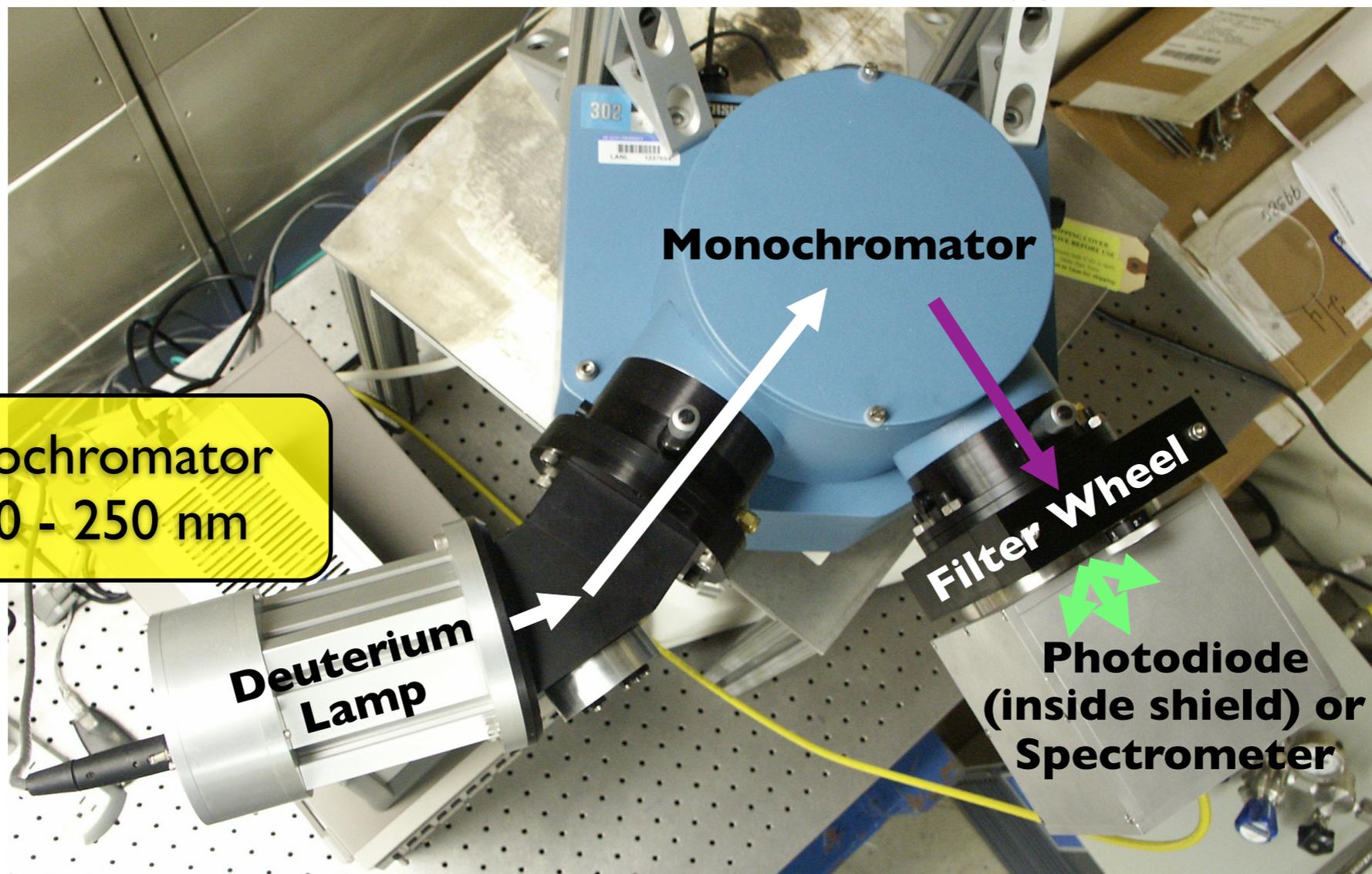
$$\epsilon(\lambda) = \frac{I_{\text{TPB}} - I_{\text{dark}}}{I_{\text{lamp}} - I_{\text{dark}}} \times g \frac{\int d\lambda' \frac{hc}{\lambda'} C(\lambda') S(\lambda - \lambda')}{\int d\lambda'' \frac{hc}{\lambda''} C(\lambda'') R(\lambda'')}$$

- Measured by us
- Measured by IRD/NIST
- Calculated from our measurements

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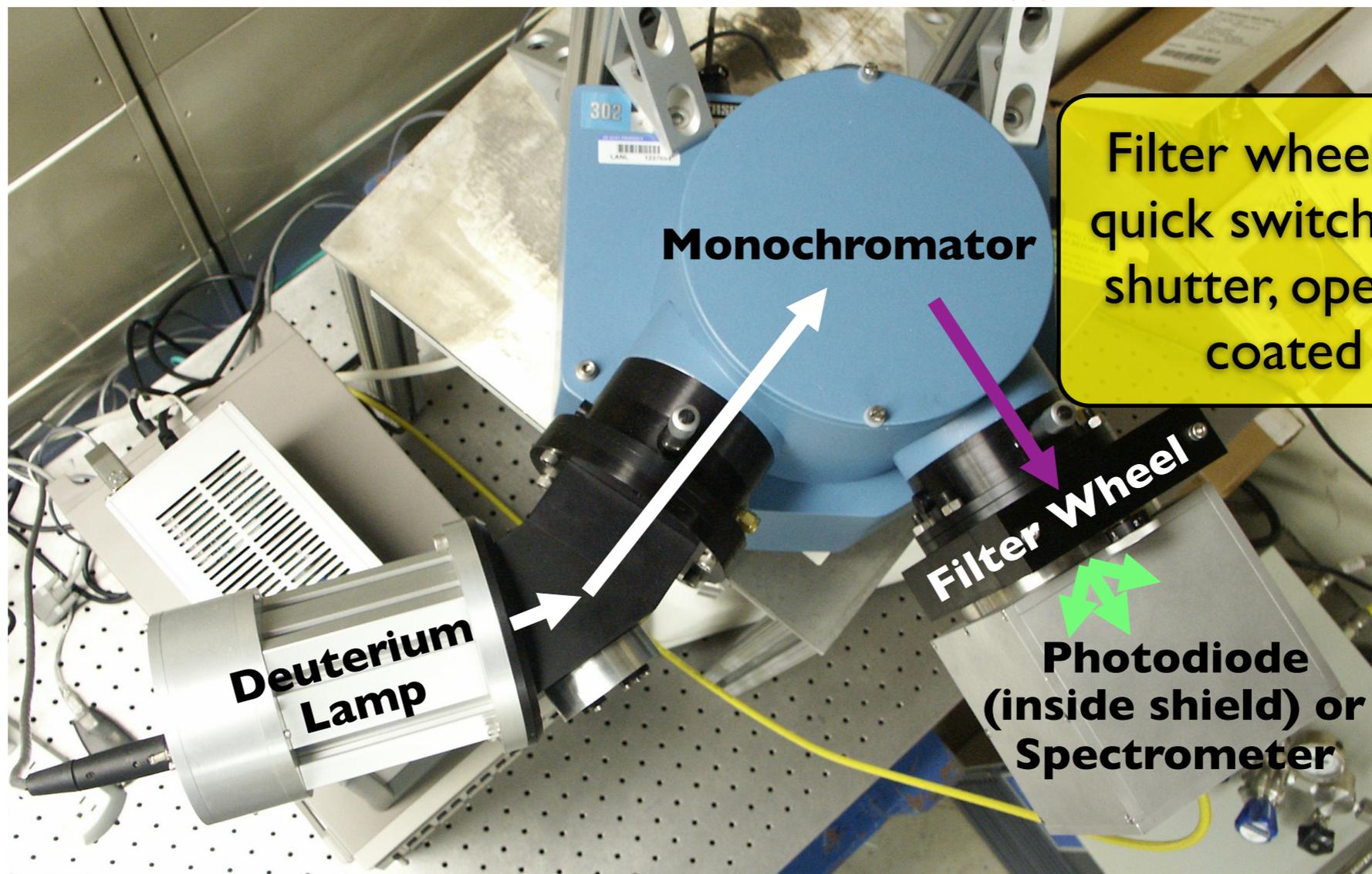
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Filter wheel allows for quick switches between shutter, open and TPB-coated acrylic

Deuterium Lamp

Monochromator

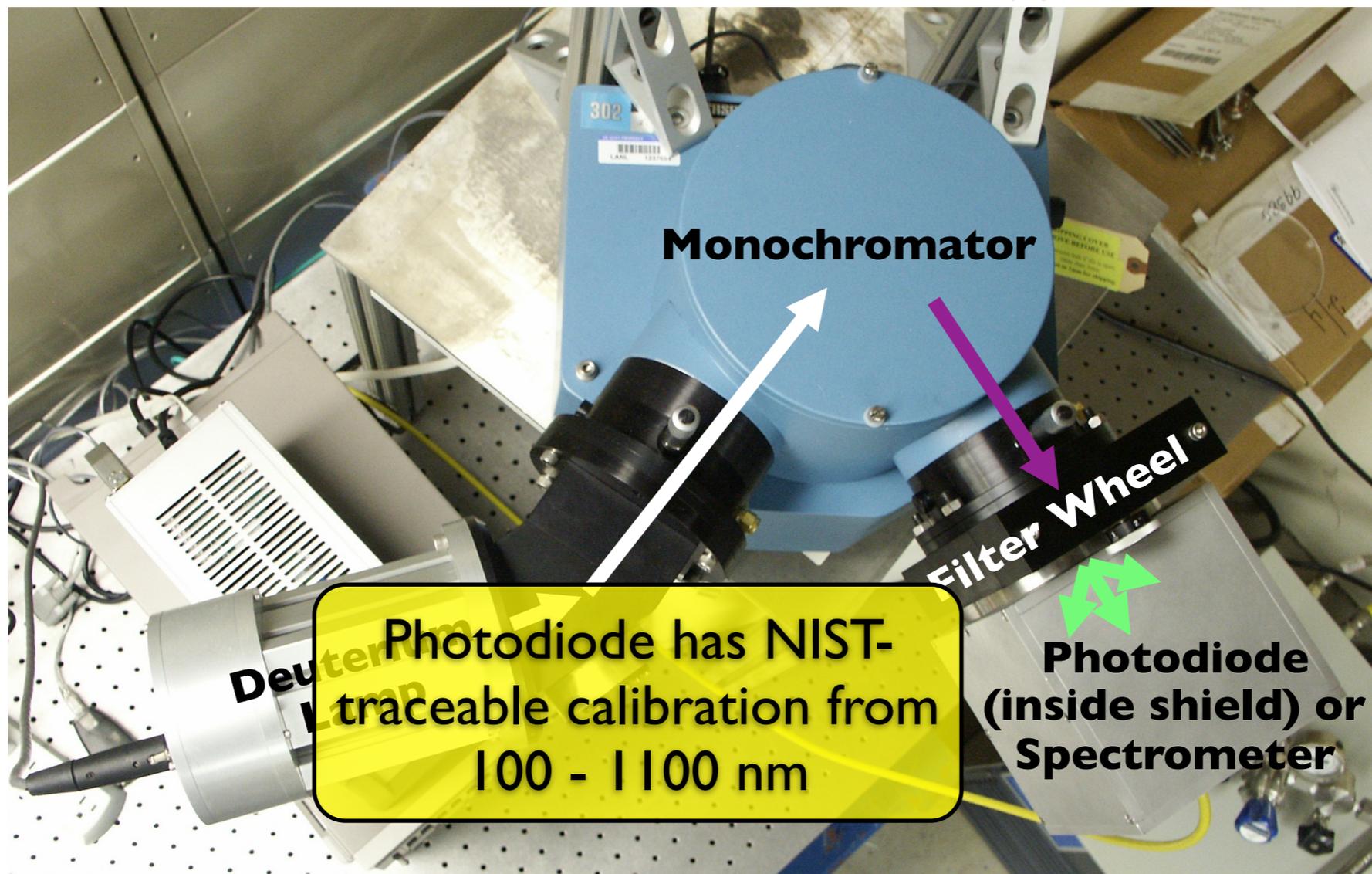
Filter Wheel

Photodiode (inside shield) or Spectrometer

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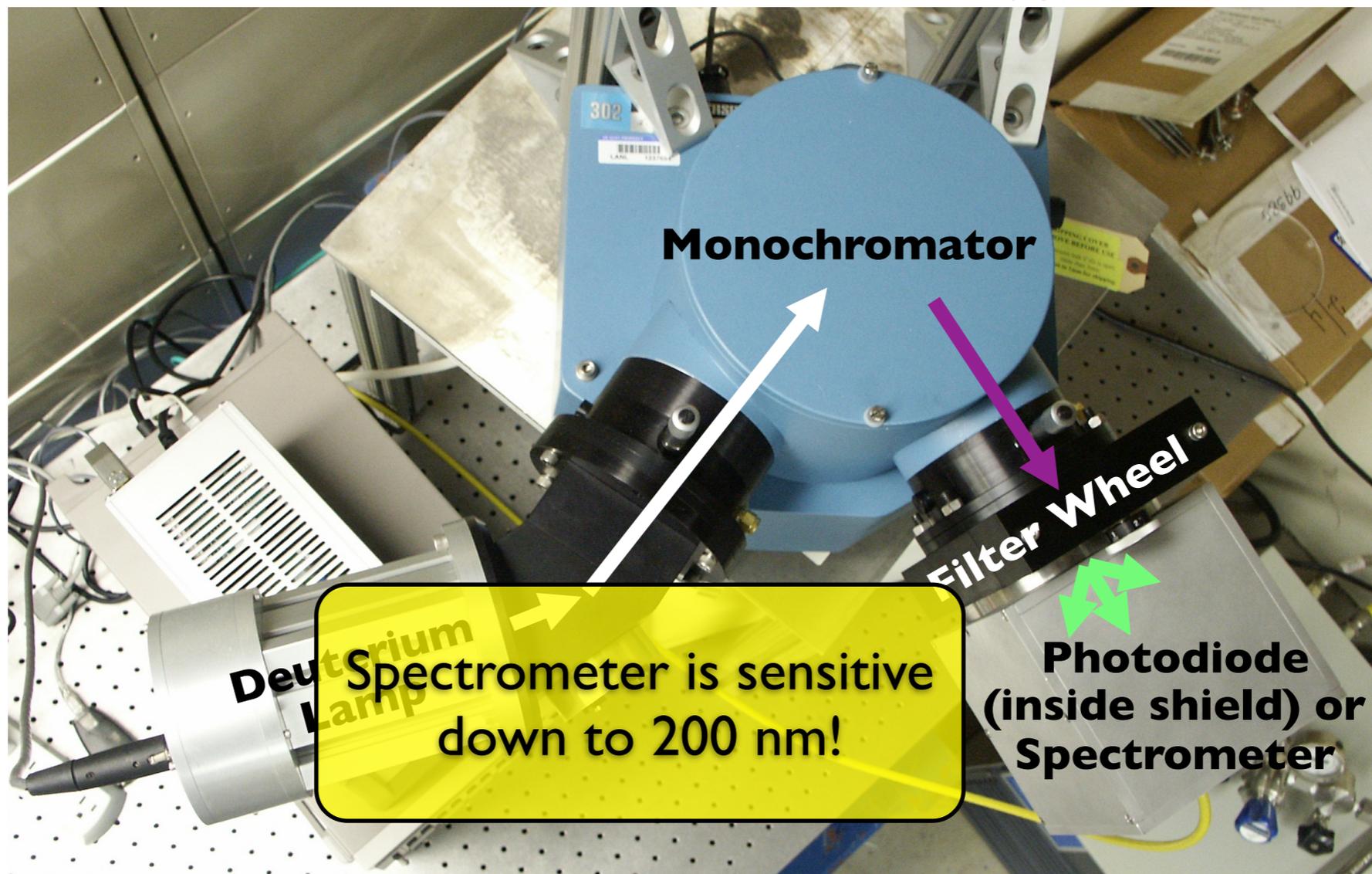
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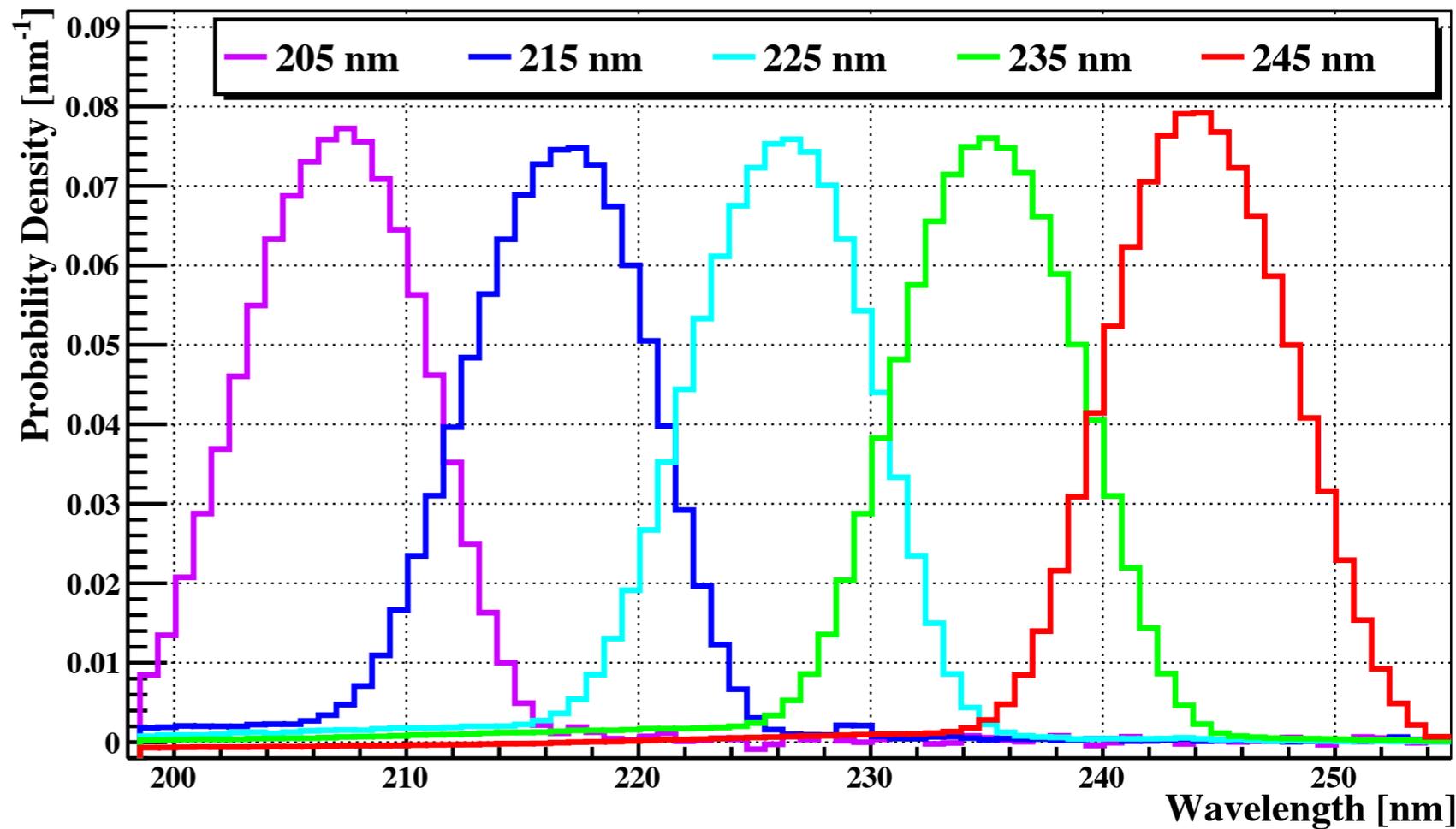
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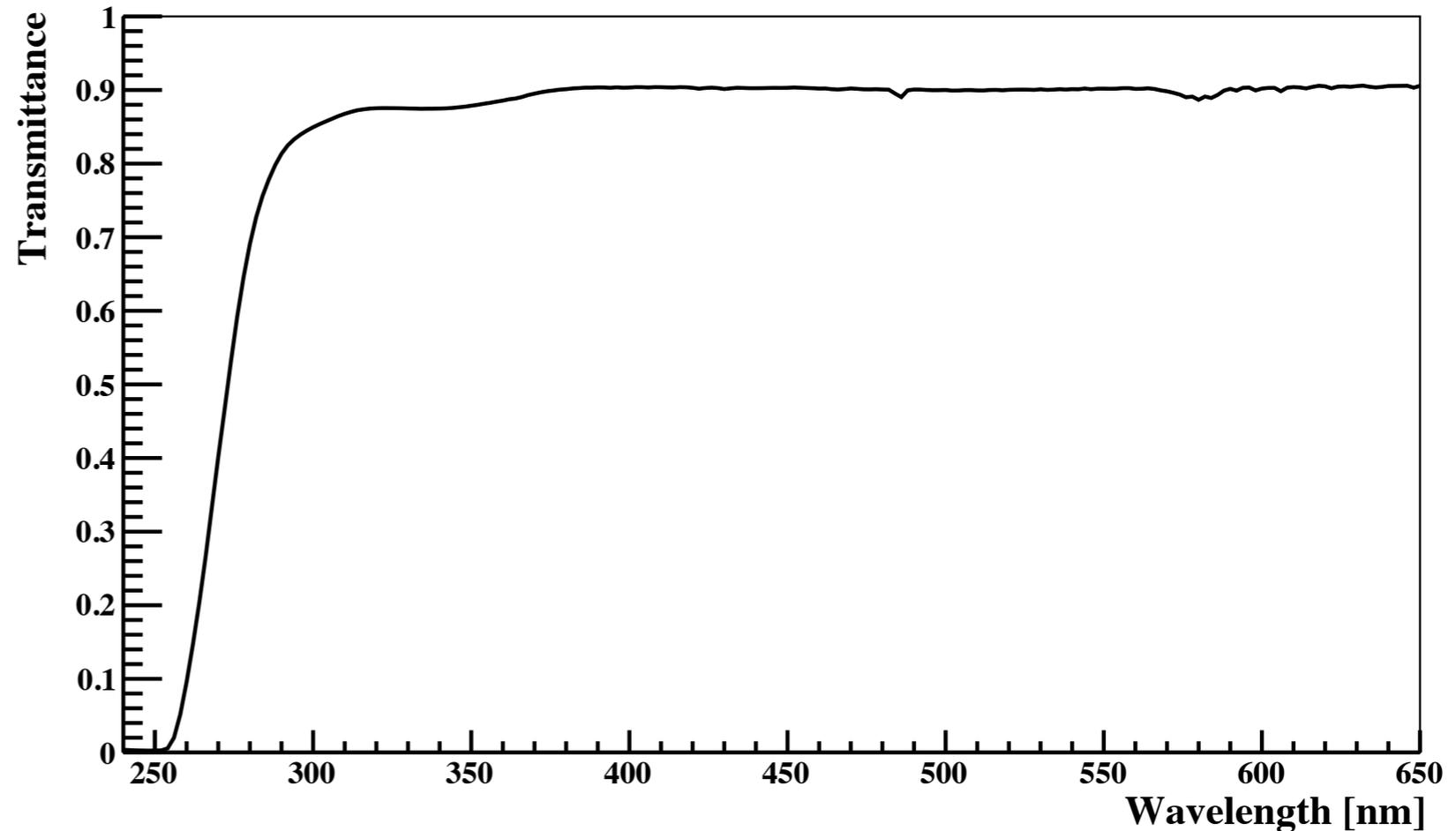
Systematic Checks

- Light source resolution: 8.5 ± 0.5 nm
- TPB film thickness: 1.5 ± 0.05 μm (thin film reflectometry)
- Acrylic (substrate) transmittance
- Optical train (lens, fibers, feedthroughs) transmittance
- Photodiode response (measured by IRD and NIST)



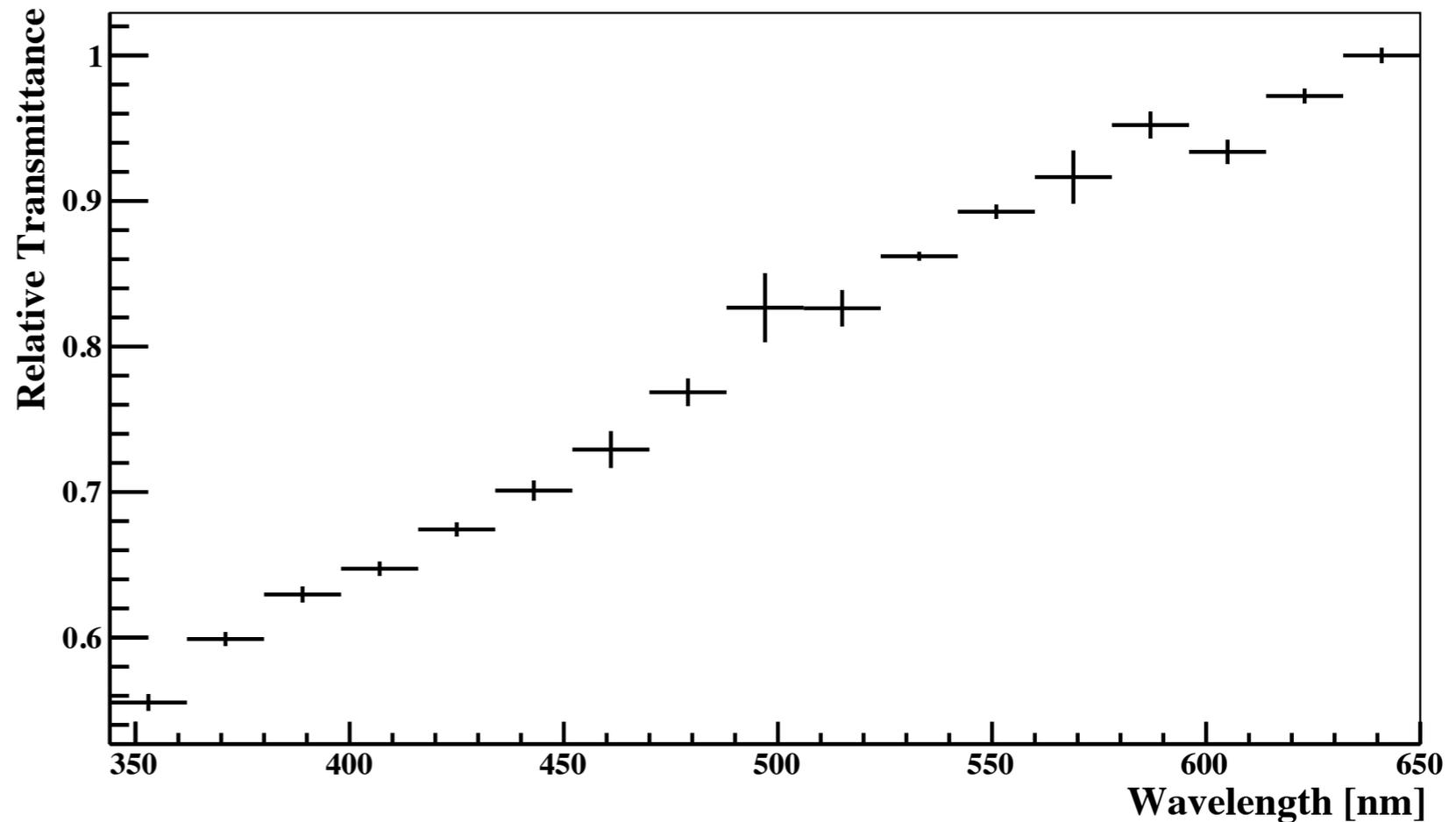
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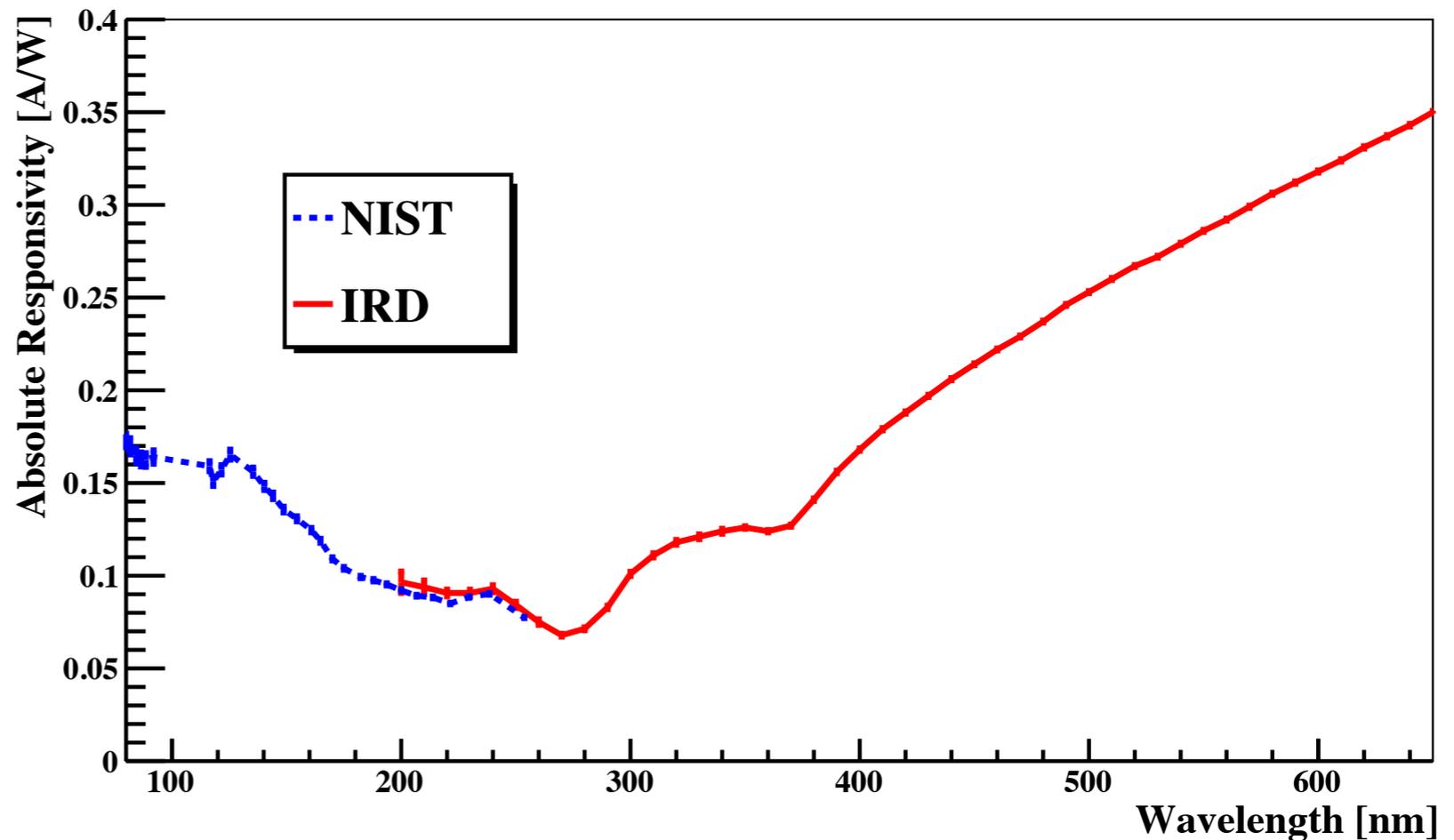
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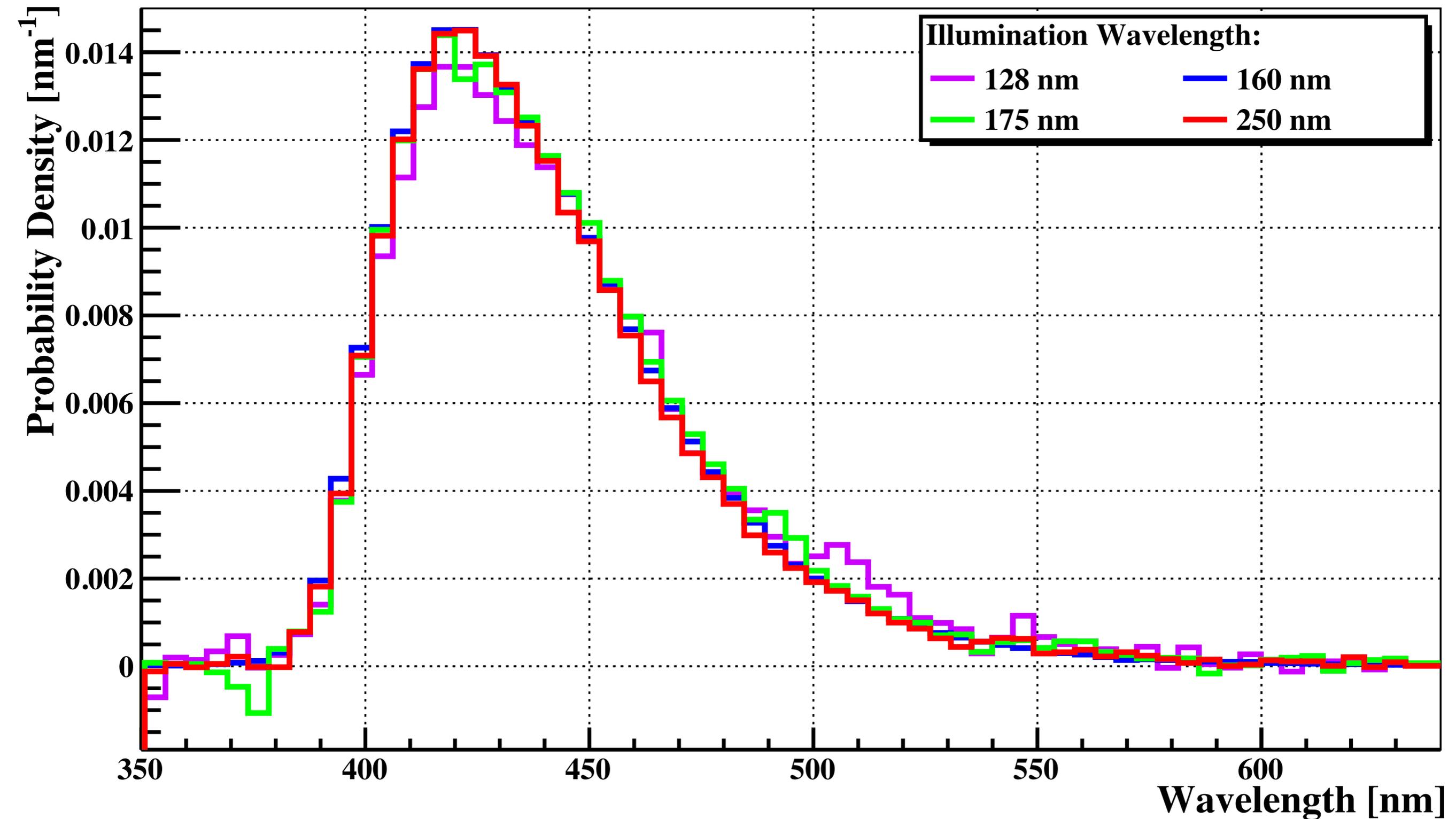


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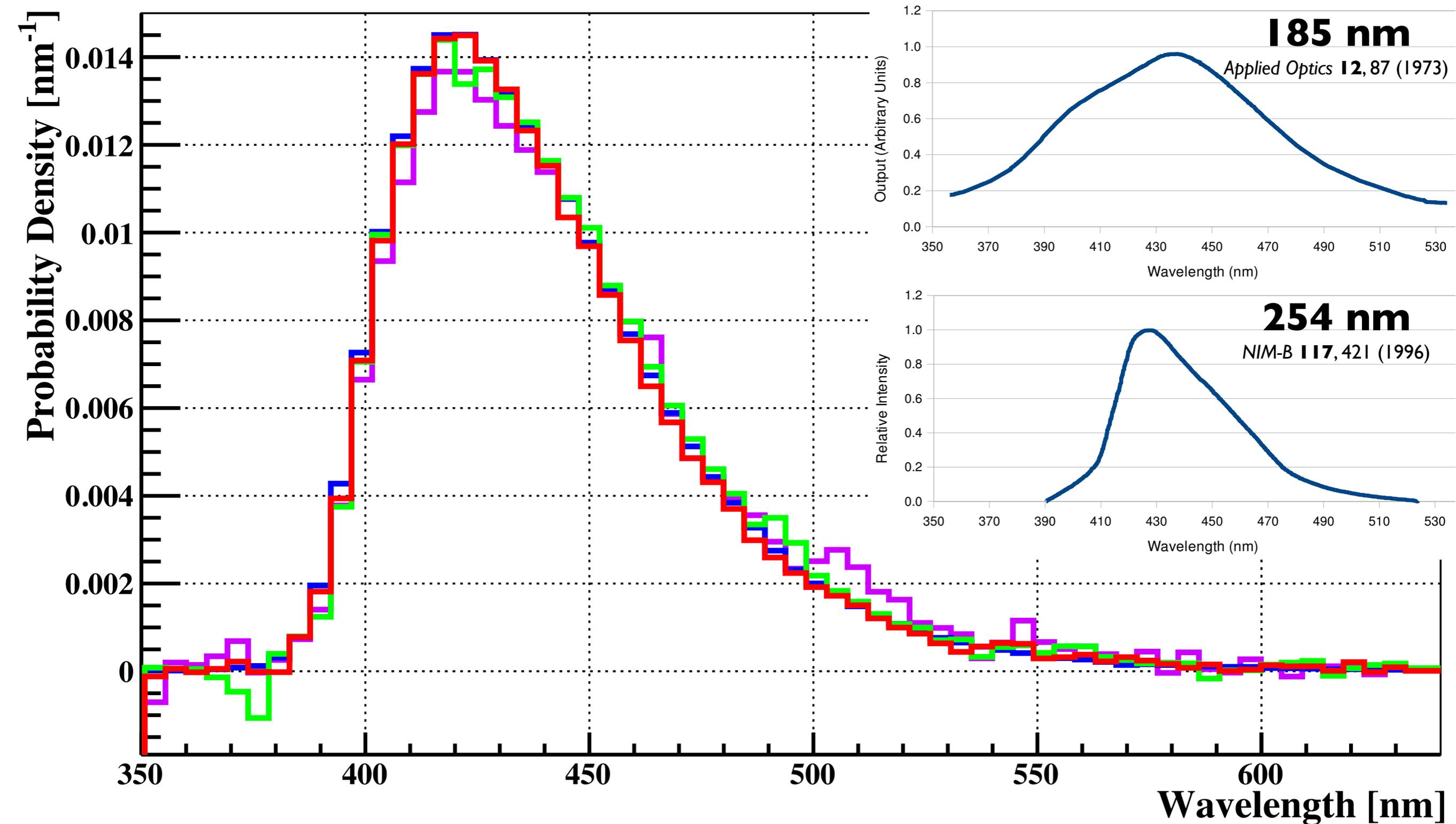


Visible Re-emission Spectrum



No strong wavelength dependence!

Visible Re-emission Spectrum

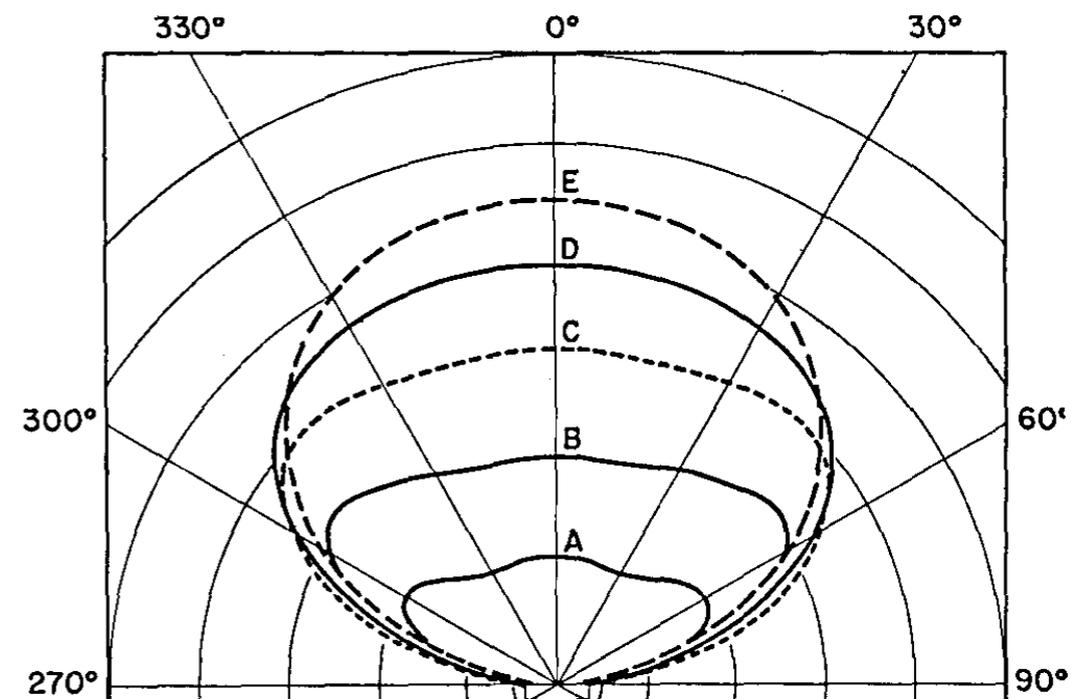


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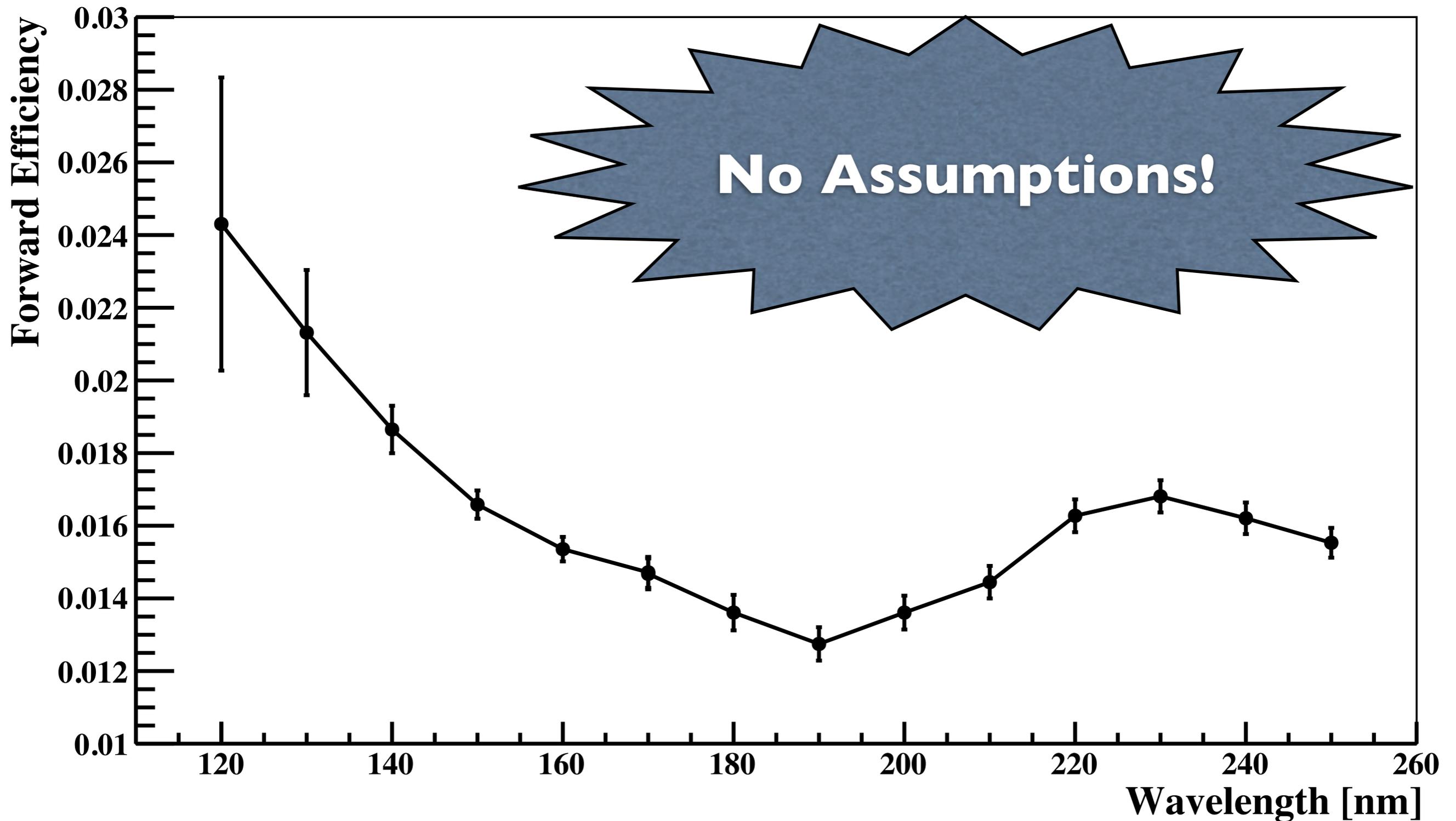
Fluorescence Efficiency

- The absolute efficiency calculation adds one more ambiguity: *angular distribution of re-emission light*
- No published measurements for TPB!
- Most naive assumption (isotropic re-emission) gives unphysically high efficiencies
- Found published angular distribution for *Sodium Salicylate* (the “other fluor” from a few slides ago)
- Follows Lambertian (cosine) distribution
- Calculated “Forward Efficiency” (re-emission at 0° , no assumptions) and total efficiency (more useful, requires Lambertian assumption)

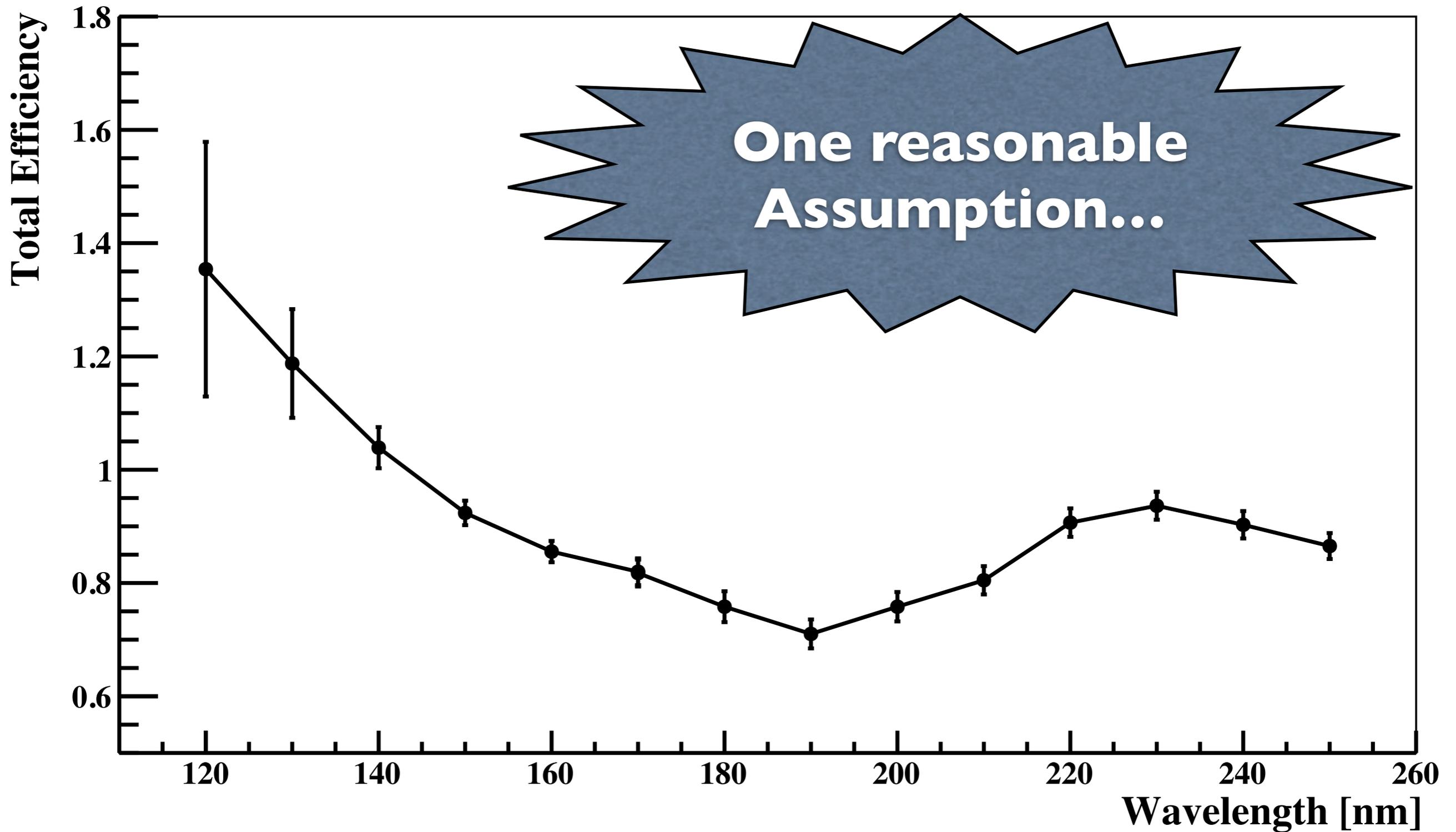
J. Opt. Soc. Am., **54**, 747 (1964), Fig. 3



Forward Fluorescence Efficiency



Total Fluorescence Efficiency

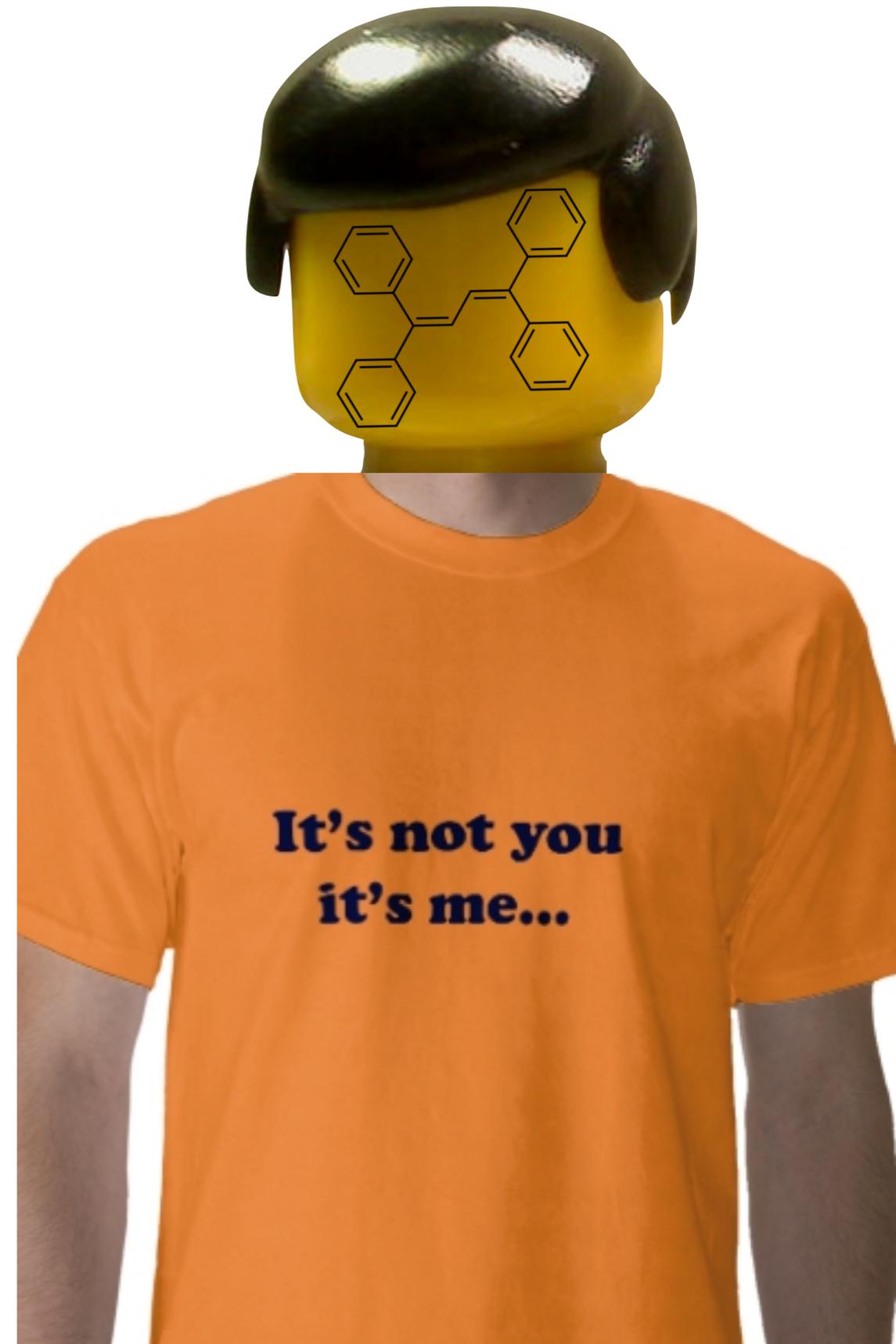


Plans for Future Work at LBL

How I would like to spend a good chunk of next year...

Other WLS Films?

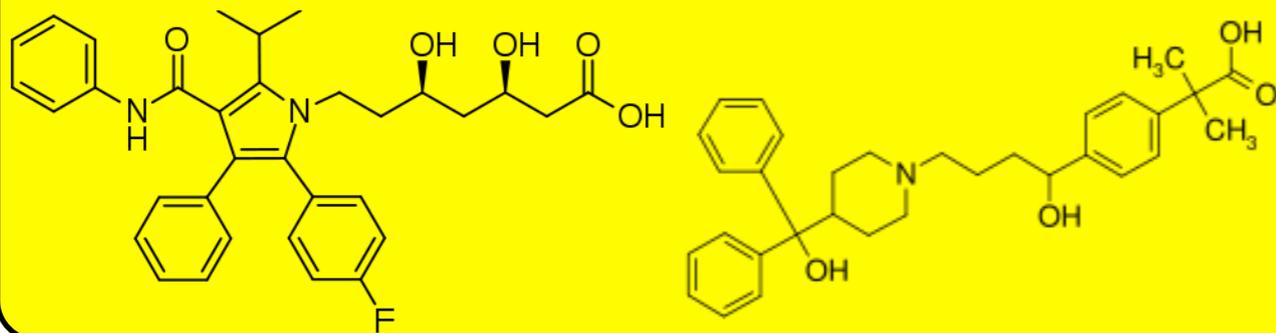
- Vacuum deposited TPB nicely, and it's been well studied. Why would we try something else?
- TPB is kind of expensive (\$113/5g)
- Vacuum deposited films are a fragile
- There are ways of making more robust TPB films
 - *so far*, at the expense of fluorescence efficiency
 - Other ways of doing solvent dilution/painting, coating paddles with embedded fibers
- There are plenty of other fish in the sea... TPH, Bis-MSB, PPO/POPOP, other organics with a few phenyl groups



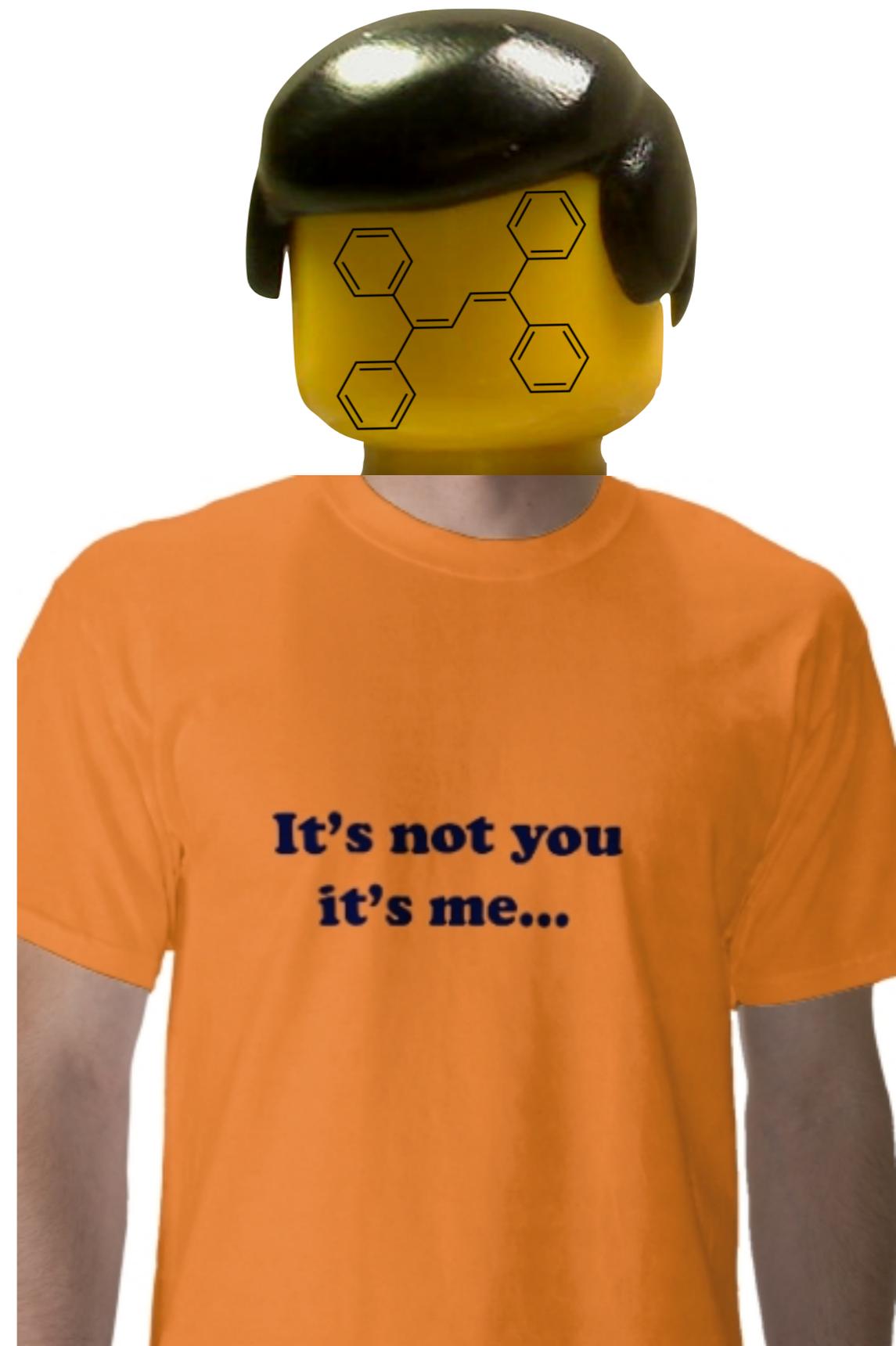
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Lipitor and Allegra also have lots of phenyl groups!

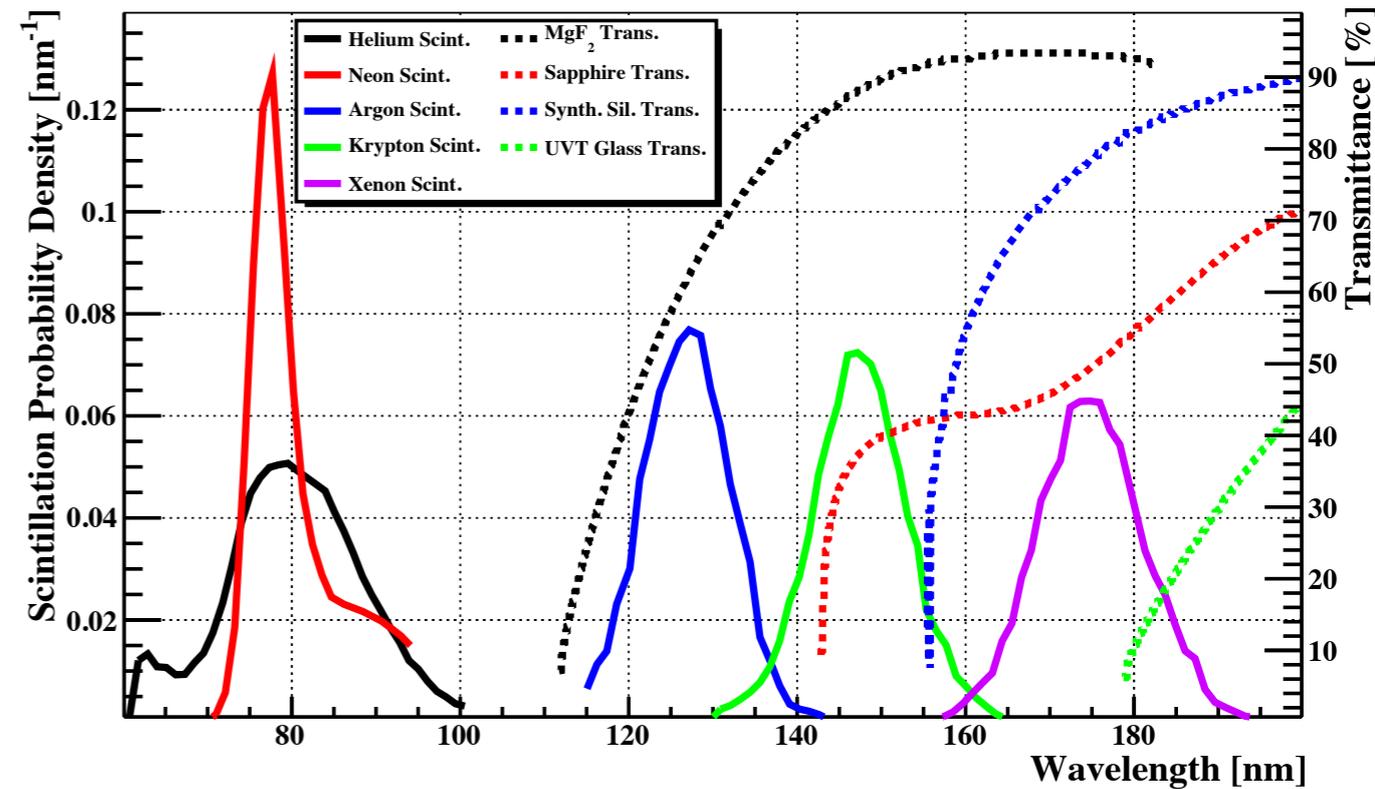


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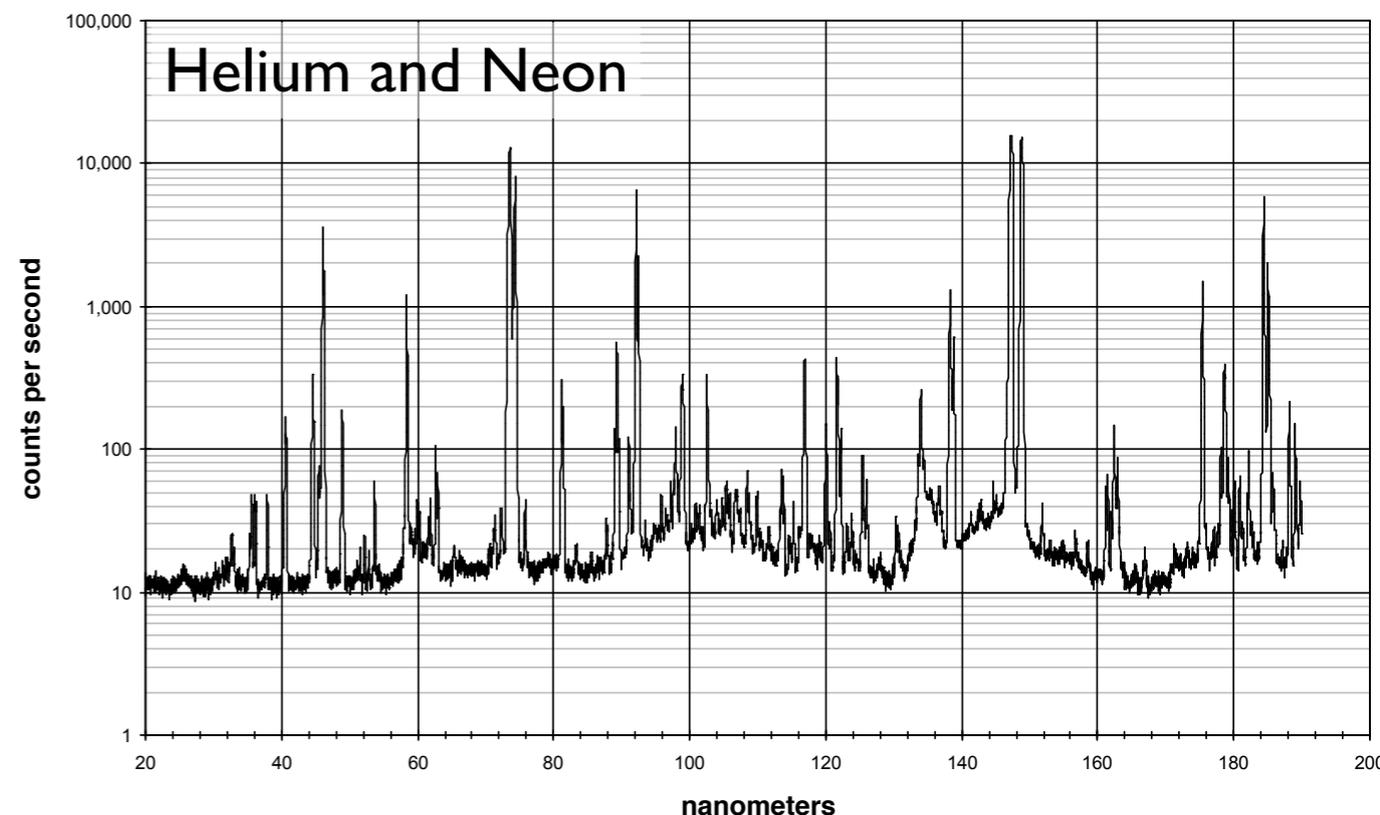
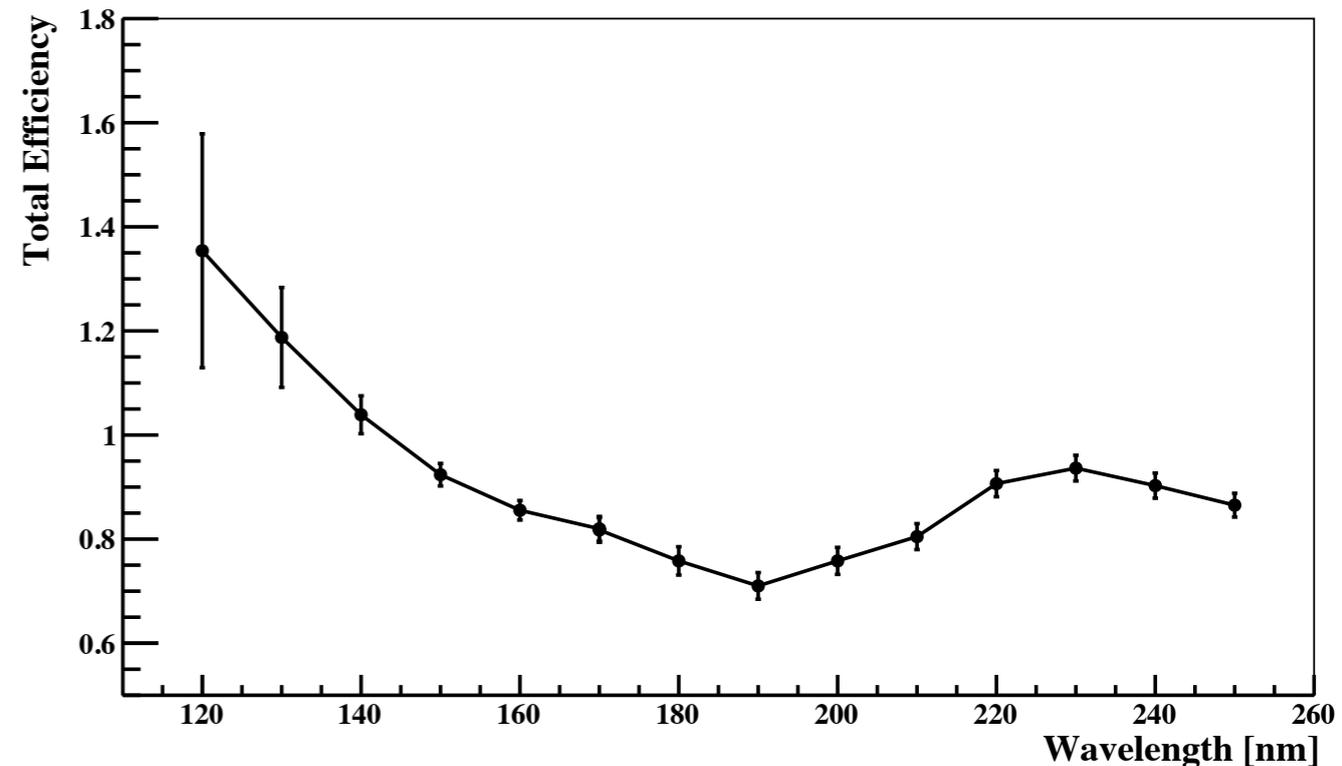
Push Fluorescence Studies Down to below 80 nm

- The MgF_2 window on our lamp limited the short wavelength range of our measurements
- This would make these measurements applicable to neon and helium
- Brighter intensity at 128 will shrink down the uncertainty at short wavelength.
- Can also change gases to get different spectra for different measurements...
- More comprehensive understanding of fluorescence behavior



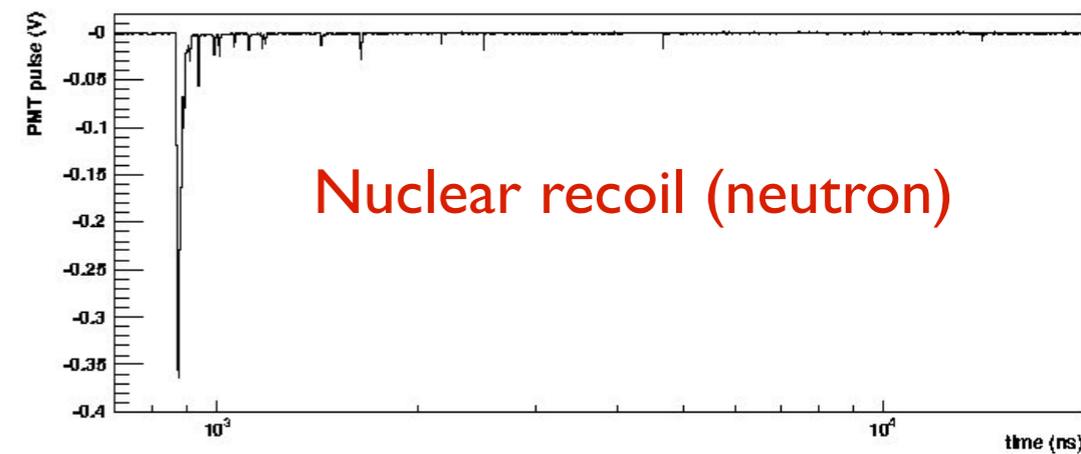
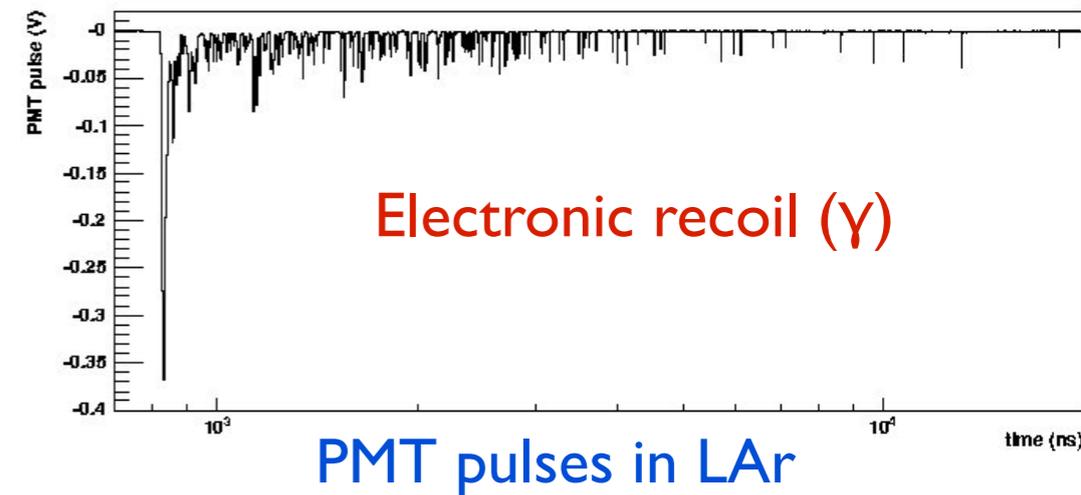
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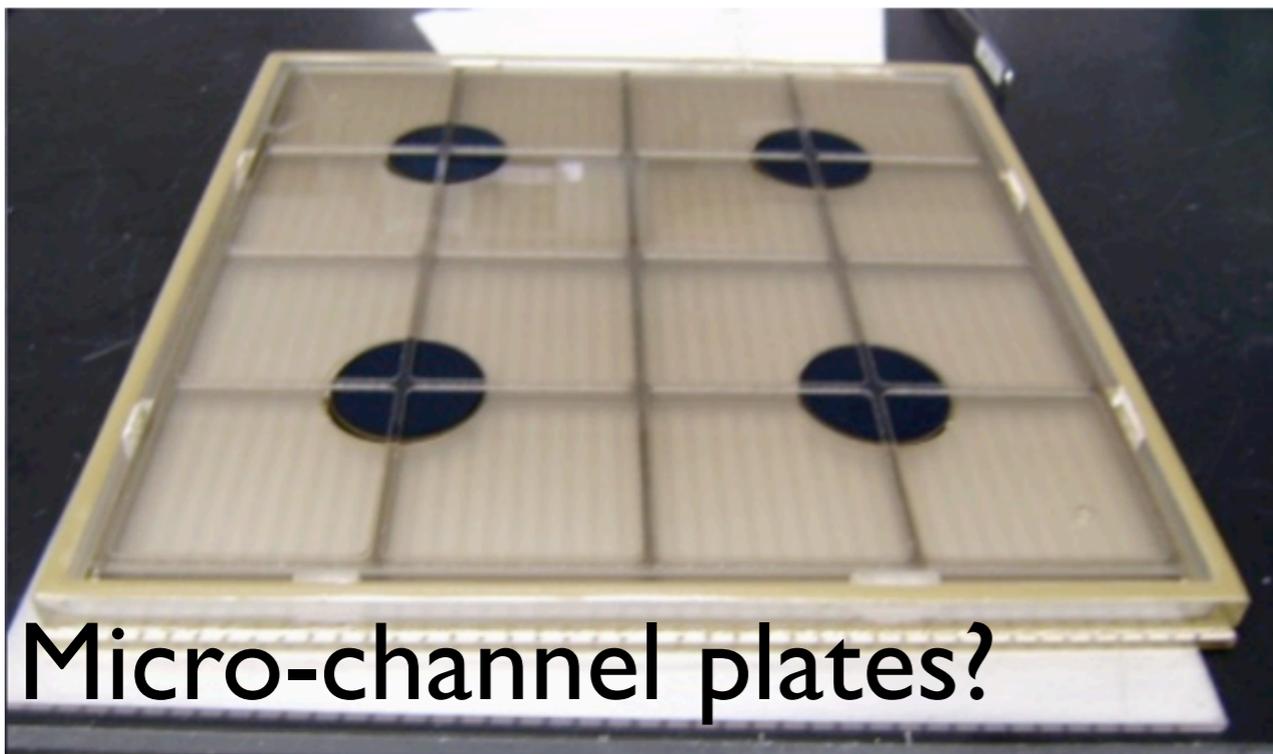
Fluorescence Time Structure

- Pulse shape analysis (singlet vs. triplet light) is one nice way of tagging nuclear and electron recoils
- How good a job can you do?
 - Are you limited by PMT response, or WLS?
 - Is there a temperature response?
- This is actually quite important if you're trying to model your PSA response for design of readout electronics, *etc.*

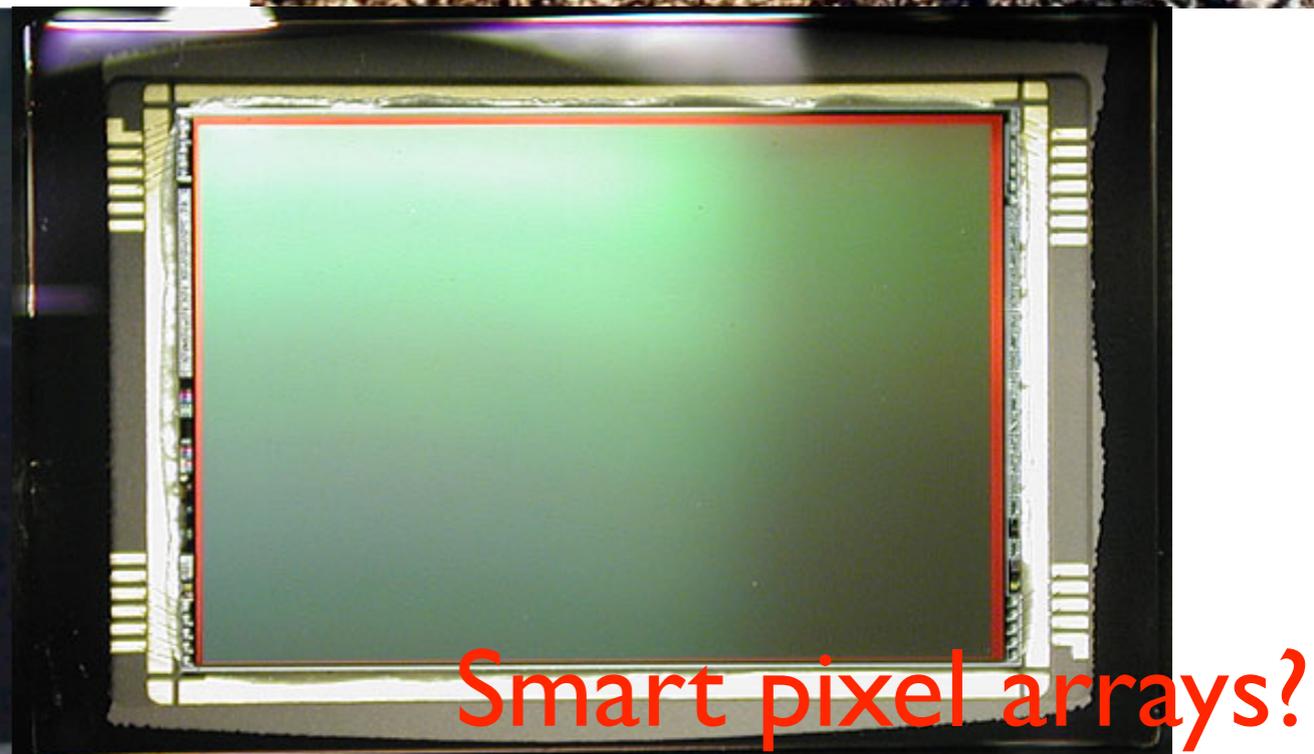


Something Completely Different?

- Wouldn't it be nice to not have to launder all of our light through a wavelength shifter?
- The alternatives clearly have problems, but maybe it's time for another look!



Micro-channel plates?



Smart pixel arrays?

Conclusions

- Detection of EUV photons is about to be *really* important to a lot of interesting particle and nuclear physics experiments
- There are some nice optical measurements and detector development projects that can dramatically influence the design and scope of these experiments
- EUV photon detection problems are *not*, in general interchangeable: right design for neutrino oscillation \neq $\beta\beta$ or dark matter
- Multi-disciplinary laboratories are well-positioned to contribute mightily to this problem!!!

Thank you for your attention!

Any questions?

